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1 **Insights into fluid transport mechanisms at White Island from analysis of coupled very**  
2 **long-period (VLP), long-period (LP) and high-frequency (HF) earthquakes**

3

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14 **Abstract:** The August 2012 to October 2013 White Island unrest sequence included 5  
15 explosive volcanic eruptions and emplacement of a small dome. These events were linked to  
16 an overall increase in SO<sub>2</sub> and H<sub>2</sub>S gas flux and RSAM seismic tremor which began in late  
17 2011. Prior to this unrest, a small swarm of 25 events was observed on 19-21 August 2011  
18 and captured on a temporary seismic array including 14 broadband sensors. Each event  
19 comprised coupled pulses having distinct high frequency (HF =2-5 Hz), long-period (LP =  
20 0.5-1.1 Hz) and very long period (VLP = 0.04-0.125 Hz or 8-25 s) earthquakes.

21 For each coupled event, we compute the source locations, origin times and related  
22 uncertainties by application of standard arrival time locations for the HF earthquakes and  
23 waveform semblance for the LP and VLP earthquakes. Results suggest that the events are  
24 centred beneath the active vent at depths generally less than 1.5 km. The HF and LP  
25 earthquakes have shallow depths (less than 1 km), while VLP have slightly deeper source  
26 locations (0.8-1.5 km). Emergent onsets for LP and VLP sources make an analysis of the  
27 absolute origin times problematic but waveform matching of VLP to LP and HF components  
28 suggests that the main VLP pulse precedes the HF and LP source processes.

29 Waveform inversion for the VLP source is consistent with the rupture of a high angle East-  
30 West oriented crack opening either in a purely tensile or shear-tensile manner. The moment  
31 of the isotropic component is estimated at  $1.2 \times 10^{12}$  Nm and the corresponding volumetric  
32 change is about 145-450 m<sup>3</sup>.

33 Results are interpreted as an upward migration of fluids which first excite the VLP from a  
34 high angle crack in the magma carapace followed by the excitation of LP and HF source  
35 processes in the overlying hydrothermal system.

36

37 Keywords: Volcano seismology; Waveform inversion; Volcanic gases; Numerical modelling

## 38 **1. Introduction**

39 The nature and character of volcanic plumbing systems may be elucidated from a range of  
40 observations including gas geochemistry (Christenson and Wood, 1993; Francis et al, 1998;  
41 Burton et al, 2000; Werner et al., 2008), seismology (Chouet et al., 1996; Neuberg et al.,  
42 1998; Kumagai et al., 2003; Bean et al., 2014), geophysics (Ingham, 2009; Legas et al., 2009)  
43 and an interpretation of the chemistry and character of the eruptive products (e.g., Houghton  
44 and Nairn, 1991; Valentine and White, 1991; Kilgour et al., 2010; Lube et al., 2014).

45 White Island volcano (Fig 1) is a frequently active composite cone with a deeply excavated  
46 central crater. It is a well-studied island volcanic system thought to include an upper to mid-  
47 crustal magma chamber (Cole et al., 2000; Werner et al., 2008), a shallow hydrothermal  
48 system (e.g., Fournier and Chardot, 2012; Chardot et al., 2015, Heap, 2015) and having  
49 pervasive magmatic degassing (Werner et al., 2008) which suggests the presence of shallow  
50 magma over periods of decades and longer. The long term degassing trend may indicate the  
51 presence of a shallow convecting magma system (e.g., Kazahiya et al., 1994; Stevenson and  
52 Blake, 1998) or alternatively to the pervasive injection and crystallisation of magma to  
53 shallow depth (e.g., Cole et al., 2000). For the former, it is uncertain if convection is possible  
54 for degassed and viscous magma. For the latter, it is uncertain if magma injection rates can  
55 yield sufficient degassing without attendant deformation associated intrusion. Evidence for  
56 deeper magma (18-22 km) at White Island is inferred from receiver functions of teleseismic  
57 waves (Park and Kim, 2014).

58 White Island volcano produces a wide range of seismic activity, including tremor and discrete  
59 LP activity (Sherburn et al., 1998) and volcano-tectonic activity (Nishi et al., 1996), with  
60 frequencies  $>1$  Hz, which are here denoted as high frequency (HF) activity to specifically

61 remove reference to rock failure processes and to generalise the event types to other possible  
62 genetic processes. This range of activity was described in detail from data collected during  
63 an earlier dense array deployment to White Island in 1992 (Nishi et al., 1996). In addition to  
64 these previously identified event types, here we note for the first time that very long-period  
65 (VLP) earthquakes (e.g., Kawakatsu et al., 1992; Neuberg et al., 1994) are also present at  
66 White Island. The prior VLPs occurred infrequently, were characterised by wave periods of  
67 ~8-25 s and were first observed with the establishment of broad-band seismic monitoring in  
68 2007. Before this period, seismicity was monitored using short-period seismic sensors and  
69 VLP activity may have occurred but remained undetected.

70 In the 14 month period from 5 August 2012 to 11 October 2013 White Island experienced 5  
71 ash and steam eruptions (Chardot et al., 2015), numerous small scale mud and ash eruptions  
72 (Jolly et al., 2016) and the extrusion of a small lava dome sometime in September-November  
73 2013 (Fig 2). The ash and steam eruptions were recorded on the WIZ seismic station which  
74 was joined by station WSRZ in April 2013. Before the eruption sequence GNS Science  
75 deployed 14 temporary broadband seismic sensors for approximately 6 months (Fig 2). The  
76 deployment period was characterised by mostly low level activity, however, twenty-five  
77 mixed frequency events including very long period (VLP), long-period (LP) and high  
78 frequency (HF) components were observed on 19-21 August 2011 (Fig 3 and 4). This paper  
79 focuses on the 19-21 August activity which, in hindsight, may have been the first precursor to  
80 the 2012-2013 unrest periods.

81 The coupled events provide a valuable opportunity to examine the source processes that may  
82 produce such seismicity in volcanic regions owed to their close temporal and spatial  
83 relationship by examination of a small sequence of coupled HF, LP and VLP events. We  
84 focus on constraining the timing and spatial relationship between the coupled events through  
85 the use of specific seismological tools. For the HF component we rely on arrival time

86 location methods (e.g., Lee et al., 1972) to obtain event locations from a well constrained  
87 upper crustal velocity model (Jolly et al., 2013). For the LP and VLP components we utilised  
88 a waveform semblance approach (e.g., Kawakatsu et al, 1999; Almendros et al., 2002) to  
89 constrain their location. The relative timing of each coupled event is examined by analysis of  
90 waveform onsets and from waveform matching techniques.

91

## 92 **2. Data**

93 Coupled HF, LP and VLP events were captured on a temporary dense array deployment of 13  
94 broad-band sensors that were operating between June and November 2011. The temporary  
95 array comprised Guralp 6T and 40T sensors and 24 bit Nanometrics Taurus digitisers which  
96 recorded continuously at a sampling rate of 100 Hz. The temporary network joined a single  
97 permanent station WIZ (see Fig 1) equipped with a Guralp 3ESP sensor with a 24 bit  
98 Quanterra digitiser that also sampled at 100 Hz. All sensors had flat response characteristics  
99 from either 30 or 60 seconds to the Nyquist frequency (50 Hz) and were processed as  
100 appropriate for the subsequent set of analysis.

101 The 6-month deployment period was remarkably quiet in terms of local seismicity, including  
102 low level tremor and very little discrete volcanic earthquake activity. Surface features at this  
103 time were also muted with no notable surface activity. The 19-21 August swarm event was  
104 initially recognised by visual inspection of the real time data on station WIZ. An STA/LTA  
105 algorithm was applied to the entire 6-month period to extract relevant low frequency events.  
106 This procedure highlighted 25 events recorded from 19-21 August and other triggered events  
107 that had mostly teleseismic origin.

108 As a first step in processing, we examined a regional event to confirm the polarity of stations  
109 and components in the temporary and permanent network. We also identified failed  
110 components (WI02, E-component) and sites with poor long period response characteristics  
111 (WI04, both horizontal components), an effect that may be associated with the installation  
112 rather than a specific site effect. Initial examination of the August 19-21 sequence revealed  
113 coupled VLP, LP and HF frequency components for each event (see Fig 3 event 16 and 17  
114 for example).

115 Analysis of the frequency characteristics of a good signal to noise example event (Fig 4A)  
116 allowed an overview of the specific spectral characteristics of the earthquake sequence with  
117 HF earthquakes having spectral bands from 2-5 Hz, LP frequencies from 0.5 to 1.1 Hz, and  
118 VLP having 0.03 to 0.125 Hz (Fig 4B). The frequency content of the HF events is  
119 conspicuous when compared to the broader spectrum associated with classical rock failure  
120 volcano-tectonic (VT) earthquakes, an observation which may reflect source process or high  
121 levels of noise at White Island. This is examined further in the discussion section. The zero-  
122 phase band passed waveforms suggest that the main VLP onset precedes the HF and LP event  
123 by a few seconds (Fig 4C), an observation that is also seen in other events of the sequence  
124 (Fig 3). This observation also is consistently seen when a strictly causal filter is applied,  
125 which implies that it is not an artefact of the zero-phase filters used in most of our analysis. In  
126 addition, a weaker VLP onset at about 30 seconds is seen in high signal to noise examples  
127 (e.g., Fig 3, events 17, 19, 20). The features in individual events are also seen from stacked  
128 unfiltered waveforms (analysis not shown) confirming that the VLP indeed precedes the  
129 LP/HF activity.

130 To understand the temporal and spatial patterns for the sequence, we utilised classic  
131 waveform semblance (Kawakatsu et al, 2000) to locate VLP (Table 1) and LP (Table 2)  
132 events and a standard arrival-time approach to locate HF sources (Table 3). We also

133 performed full waveform inversion for the moment-tensor the and the time-history of the  
134 VLP source. All the outlined methods rely on the knowledge of velocity structure which we  
135 discuss next.

136

### 137 **3. Shallow velocity structure at White Island**

138 The shallow velocity structure at White Island was perviously constrained from an active  
139 source experiment conducted in November 2011 at the end of the dense array deployment  
140 (Jolly et. al., 2013). This work found that the unconsolidated tephra within the shallow crater  
141 floor had low velocities of approximately 1.2 km/s while the crater walls had velocities of 2.2  
142 km/s. Ray paths from the active sources only propagated to depths of a few hundred m below  
143 the surface however and deeper velocities were not sampled. To verify the active source  
144 results and sample deeper portions of the volcanic system, we applied a noise interferometry  
145 approach using the temporary dense array data. We analysed a linear array of temporary  
146 seismometers extending in a northwest-southeast direction paralleling the valley floor. These  
147 data were supplemented by a broadband station (WI05) on the north-western flank of the  
148 edifice to derive a shear-wave velocity model of the upper ~1.0 km of the crater area of  
149 White Island. The dense crater floor array had a length of about 0.8 km (see Jolly et al., 2013,  
150 their Fig 1). The linear aperture is further increased to about 1.5 km with the addition of the  
151 north western station (WI05). We use noise interferometry to generate dispersion curves  
152 between about 0.5 and 5 Hz (as in Stehly et al, 2009; Fry et al, 2010). This wide bandwidth of  
153 dispersion observations is obtained by generating a single composite dispersion curve,  
154 selecting the frequency range contribution of each station pair according to its inter-station  
155 distance. We aim to investigate the uppermost km of the crater floor so we first decimated  
156 the data to 10 Hz. Other pre-processing steps include removing the instrument response and

157 the mean from the seismic data, and whitening the signal to remove the spectral peaks of the  
158 spectrum to satisfy the requirement of equipartition for ambient-noise analysis. The latter is  
159 necessary to assume that the cross-correlation function equals the Green's function between  
160 two stations.

161 We performed a time-domain cross correlation of station pair per day and stacked the data  
162 over the whole period, which is equivalent to combining the effects of sources located at  
163 different azimuths towards equipartition. A group velocity was then computed for each  
164 station pair, after multiple filtering to isolate the fundamental mode. Rayleigh waves are  
165 dispersive, so their velocity depends on the frequency. We therefore extracted the phase  
166 velocity for each station pair and each frequency and obtained a dispersion curve with a  
167 standard deviation by summation. Since the Rayleigh wave velocity is a function of  $V_p$ ,  $V_s$ ,  
168 density and the layer thickness, this dispersion curve was finally inverted to obtain a 1-D  $V_p$ ,  
169  $V_s$  and density profiles. We invert the resulting dispersion curves with the geopsy (Wathelet,  
170 2008) software to obtain a layered 1D  $V_s$  model. Our results allow us to generate a multi-  
171 resolution model with ~5 m vertical resolution in the upper 20 m and ~200 m resolution at 1  
172 km depth. We define interfaces at ~4 m, ~80 m, and ~1100 m. Below 1100 m, we define  $V_s$   
173 ~ 1.225 km/s and  $V_p$  ~ 3.25 km/s. Results suggest that on average, the crust beneath the  
174 volcano experiences a strong shallow velocity change at about 90 m in the central crater  
175 where  $V_p$  ~ 1.2 km/s underlain by a ~1 km layer with  $V_p$  ~ 2.4 km/s. These results are  
176 consistent with velocity estimates obtained from the active source experiment (Jolly et al.,  
177 2012) with low velocities ascribed to unconsolidated tephra and higher velocities ascribed to  
178 the underlying consolidated volcanic rocks. Interestingly, the  $V_p/V_s$  ratio in the region is  
179 relatively high at about 1.95, suggesting the presence of at least a small amount of volatiles  
180 and hot rock in the uppermost km (Fig 5).

181 Because the active source and dispersion curve results match closely, with only ~0.2 km/s  
 182 difference for the upper km of the analysis, we completed our analysis of the swarm locations  
 183 using a uniform velocity of both 2.2 km/s and 2.4 km/s as reasonable range of velocity model  
 184 uncertainties. Results of the location analysis and the reasonable alternative models are  
 185 discussed next.

186

## 187 **4. Source locations**

### 188 **4.1 Waveform semblance**

189 We used the waveform semblance approach as detailed in Kawakatsu et al. (2000), and  
 190 subsequently utilised by Almendros et al. (2002) and Almendros and Chouet (2013) for the  
 191 analysis of LP and VLP components which are in turn derived from the pioneering work of  
 192 Neidel and Tarner (1971). Coherency amongst seismograms can generally be measured  
 193 using:

$$194 \quad S = \frac{\sum_{j=1}^L \left( \sum_{i=1}^N u_{i,j(i)} \right)^2}{N \sum_{j=1}^L \sum_{i=1}^N u_{i,j(i)}^2}$$

195 , (1)

196 where  $L$  is the length of the seismogram,  $N$  is the number of stations and  $u_{i,j(i)} \equiv u_i(t_i + j\Delta t)$  is a  
 197 seismogram of the  $i$ th station at the  $j(i)$ th time sample at the start time  $t_i$ .

198 Waveform coherence implies that the stationary isotropically radiating point-source imparts  
 199 seismic energy to all stations as compressive body waves that may be rectified into a linear  
 200 particle motion (Kawakatsu et. al, 2000). In the radial component (the direction of a station  
 201 from the source) the seismogram particle motion is given by  $u_{i,j(i)} \equiv R_{i,j(i)}$ . Different stations

202 have different source station travel-times and adjustments for these time delays ( $t_i$ ) should  
 203 increase the radial components waveform coherence. There may also be mutually  
 204 perpendicular components  $H_{i,j(i)}$  and  $V_{i,j(i)}$ , which for an isotropic radiating source having well  
 205 behaved ray paths, should have negligible contributions. The waveform semblance may be  
 206 incorporated to emphasise the radial excitation from the isotropic source and penalise  
 207 departures from rectilinear behaviour via the following equation (Kawakatsu et al., 2000):

$$208 \quad S_3 = \frac{1}{D} \sum_{j=1}^L \left\{ \left( \sum_{i=1}^N R_{i,j(i)} \right)^2 - N \left( \sum_{i=1}^N V_{i,j(i)}^2 \right) - N \left( \sum_{i=1}^N H_{i,j(i)}^2 \right) \right\},$$

209

210

(2)

211 where  $V$  is the component in the direction perpendicular to  $R$  within the vertical plane which  
 212 contains both source and receiver, and  $H$  is in the horizontal component perpendicular to both  
 213  $R$  and  $V$  ( $V$  is generally not vertical).  $L$  is the number of time samples and  $N$  is the number of  
 214 stations. We take the scaling factor  $D$  as:

215

$$216 \quad D = N \sum_{j=1}^L \sum_{i=1}^N R_{i,j(i)}^2.$$

217

(3)

218  $S_3$  measures the rectilinearity of particle motion pointing via the signal coherency in the radial  
 219 direction. As in Kawakatsu et al., (2000) we normalise each seismogram so that the RMS  
 220 amplitude of each signal at each station becomes unity:

221

$$RMS_i^2 = \frac{1}{L} \sum_{j=1}^L (R_{i,j(i)}^2 + V_{i,j(i)}^2 + H_{i,j(i)}^2) = 1,$$

(4)

which removes amplitude information from the original seismograms giving equal weight to each station during the coherence computation.

In practice, for each event, we compute the waveform semblance within a grid search of the target volume. We first compute a search over a 3 km grid centred on the crater lake and from sea level to 3 km below sea level at 100 m spacing. In our case, we search to a depth of 3 km as this would be just beyond the lower limit of resolution given the aperture of the dense array deployment. At each point we compute the source to station distance, azimuth and incidence angle and rotate the  $N$ ,  $E$ ,  $Z$  seismograms for each station into  $R$ ,  $H$  and  $V$  components having an appropriate travel-time-delay based on the independently determined velocity model. Results are normalised using Eq (4) and the waveform semblance is computed with Eq. (2). For this formulation, the waveform semblance varies between -1 and 1, with 1 indicating perfect coherence without penalty from the  $H$  and  $V$  components. After determining the best fit semblance in the coarse grid search, we re-computed the semblance within a 20 m grid spacing within a 200x200x200 m search volume centred on the best fit coarse solution. Before proceeding to the location of LP and VLP sources using the semblance method, we first show the performance and error analysis using synthetic data.

240

## 4.2 Synthetic test of waveform semblance

To obtain an improved understanding of the limitations and accuracy of the method, we compute synthetic seismograms and then determined their source location in a blind test

244 using the waveform semblance technique described above. The synthetic seismograms are  
245 computed with the open source spectral-element method (SEM) code EFISPEC3D  
246 (<http://efispec.free.fr>; De Martin, 2011). The numerical domain is based on a  $10 \times 10 \times 7 \text{ km}^3$   
247 rectangular cuboid (boundary conditions: free surface at the top, absorbing boundaries at the  
248 other 5 surfaces). The surface is then altered to reflect the local topography of White Island,  
249 using a combination of a Digital Elevation Model (DEM) for the island and a bathymetry  
250 model for the surrounding seabed (Prasetya, and Wang, 2011). Water can currently not be  
251 accounted for in EFISPEC3D, but is expected to have minor effects on the waveforms  
252 computed for this test. An example of the computational setup is shown in Fig 6.

253 The receivers are located at the free surface, at positions corresponding to the receivers in the  
254 field experiment. The synthetic sources are located at a horizontal position (E 2880132 m; N  
255 6400282 m in depths of 1.0 km and 2.0 km below sea level, respectively. For the source  
256 mechanism we use explosive sources, horizontal tensile cracks and steep oblique tensile  
257 cracks striking at  $25^\circ$  from North and dipping at  $75^\circ$ . As the source time function (STF) we  
258 use a 1 Hz low-pass filtered VLP displacement waveform from an event recorded at WIZ  
259 station. This is justified by the very long wavelength for the waveforms which are dominated  
260 by near to mid-field terms; the intermediate-field displacement term is directly proportional  
261 to the time-history in the source, and the near-field term, for a small P to S difference  
262 (relatively to the dominant period of the waveform) has the same shape. Moreover, the VLP  
263 signals are strikingly similar between different stations, so it is fully justified to use the WIZ  
264 station, which has good signal-to-noise characteristics. Two different velocity models are  
265 used: (i) homogeneous ( $V_P = 2.3 \text{ km/s}$ ,  $V_S = 1.328 \text{ km/s}$ ,  $\rho = 2160 \text{ kg/m}^3$ ) and (ii) layered  
266 model described in Section 3.

267 For the blind test, we generate seismograms for 8 different combinations of velocity models,  
268 mechanisms and depths, listed in Table 4. Four of the synthetic examples include explosive  
269 sources, two examples model horizontal tensile cracks, and 2 models have oblique tensile  
270 cracks. In addition, two of the examples include a superposition of typical White Island  
271 background noise.

272 An example of relocation using the semblance approach from Section 4.1, is shown in Fig 7.  
273 The relocation of a simple isotropic source is within 100 m of the synthetic computation with  
274 differences possibly owed to minor uncertainties in the DEM, and topographic effects. The  
275 accurate recovery of the isotropic source then provides a basis for an examination of  
276 uncertainties associated with the introduction of noise, velocity structure and the departure  
277 from isotropic source geometries. Alternate input velocities (Section 3) were found to  
278 produce only minor changes in the retrieved semblance locations of less than 100 m.

279 Results of the 8 cases show the following general trends: 1) explosive sources were well  
280 recovered (error less than 200 m) by the waveform semblance, 2) horizontal tensile cracks  
281 tend to yield consistently shallow solutions and the introduction of unknown layered velocity  
282 structure tends to increase the effect, 3) oblique high angle tensile cracks tend to produce  
283 poor semblance values and significant departures from the true source location ( $> 200$  m  
284 lateral mis-location and  $\sim 100$  m depth mis-location) and the location misfits are larger ( $\sim 1000$   
285 m lateral mis-location) if the noise levels are high. 4) noise levels generally increase the  
286 misfit between true location and computed location and decrease semblance: a result that is  
287 also apparent in the natural data (Section 4.3). In general, these tests suggest that while near  
288 vertical crack geometries yield very high misfit values with the known synthetic location,  
289 these effects can be identified because the semblance will be low for a dense seismic  
290 network. We note however that sparse networks may lead to erroneously high semblance  
291 values due to undersampling of the non-uniform source focal sphere. In general, we find that

292 the location error increases for the sources with strongly non-isotropic radiation pattern  
293 across the network (such as a vertical crack).

294

### 295 **4.3 Location of LP and VLP sources**

296 After confirming the accuracy of the method using a variety of synthetic sources, we next  
297 applied the results to each of the 25 events for zero-phase offset filtered seismograms (Fig. 4)  
298 for both LP and VLP source components. As noted in Section 2, the onset of the VLP occurs  
299 ~30 seconds prior to the main VLP/LP and HF pulses. In our case, the semblance analysis is  
300 completed on the band pass filtered main phase and not the preceding low amplitude VLP  
301 signal. An example of the particle motion results at the best fit location for event 17 is  
302 shown for a VLP and LP event (Fig 8, Fig 9). This event was selected because it had good  
303 signal to noise and VLP semblance but was not the ‘best’ example (as measured by  
304 semblance), hence more representative of a good example event.

305 The map view particle motions (Fig 8A) for the VLP event show that a source located south  
306 of the summit is reasonable, but there are notable and systematic polarisation deflections  
307 from the best fit source position. Note specifically in this example the apparent deflection of  
308 WI08 and WI09 which contributes to a location south of the central crater. This is possibly  
309 owed to a significant contribution of a non-isotropic source mechanism and mixing of P- and  
310 S-phases (for details about source mechanism see Section 5). The associated depth-section  
311 particle motions (Fig 8B), where each particle motion is rotated from its azimuth projection,  
312 show that stations closer to the source have high incidence angles and further stations have  
313 shallower incidence. This result is expected for a highly coherent semblance result from a  
314 mostly uniform source.

315 In contrast, the LP example for the same coupled event 17 (Fig 9) yields significant  
316 discrepancies between particle motions across the network. In this case, subsets of stations  
317 have very different particle motion results and the weakly coherent waveform semblance  
318 estimate is very low (see negative semblance values Table 2).

319 It has been standard practice to compute these error bounds based on a semblance error limit  
320 within the search volume (Kawakatsu, 2000, Almendros 2002, Almendros 2003). Almendros  
321 (2003) discusses a criterion semblance level  $S_L = (1 - \delta S) S_{MAX}$ , based on the signal-to-noise  
322 ratio ( $SNR$ ) which may be used to establish the error bounds ( $\delta S$ ) for a semblance location.  
323 In that case, the error bounds were determined empirically from synthetic data including a  
324 known  $SNR$ .

$$325 \quad \delta S = 0.062 SNR^{-1.54}$$

$$326 \quad (5)$$

327 In our case, the semblance is noted to vary strongly at low  $SNR$  and stabilise for  $SNR > 3$   
328 (Fig 10). The empirical relation yields error limits of about 1% in our case which may not be  
329 appropriate for White Island. Instead, for VLP source locations we used an error limit of 2%  
330 of maximum semblance, which yield lateral errors about the same as the mis-locations found  
331 in synthetic tests for explosions and horizontal tensile cracks. For LP events, having  
332 consistently lower semblance values, we used a semblance error limit of 50% of maximum  
333 semblance, which yielded lateral errors similar to the vertical crack synthetics. This approach  
334 should offer reasonable first order estimates of the location uncertainty.

335 Next we examine the 25 events using the same semblance approach for individual events.  
336 From this analysis, we obtain the range of locations and associated error bounds for the 25  
337 coupled events (Fig 11). For the VLP events, we find that locations are distributed over the

338 southern crater margin and at an average depth of  $\sim 1$  km. Formal errors from the analysis of  
339 synthetic data yield lateral errors of  $\sim \pm 150$  m and depth errors of  $\sim 200$  m. However, given  
340 the uncertainties in velocity structure and possible departures from the point source  
341 assumption, it is possible that VLP sources actually originate from the same source position.  
342 The distribution of locations is somewhat larger than our established error bounds, however,  
343 with lateral errors of about 800 m E-W, 500 m N-S and 500 m in depth. We note that the LP  
344 sources for the same earthquake sequence have very different location distribution and error  
345 bounds. LP events tend to be distributed over the northern margin of the crater (Fig 11)  
346 having formal errors of about  $\pm 300$  m laterally and  $\sim 1000$  m vertically. Shallower LP  
347 events have unconstrained errors whose formal error limits are beyond the search grid. The  
348 distribution of earthquake locations is more tightly clustered laterally (within  $\sim 300$  m) but has  
349 a wide range of depths (from the surface to 900 m depth). Semblance locations for VLP and  
350 LP events are summarized in Tables 1 and 2. We note that particle motions may be  
351 significantly distorted due to topographic gradients that may be corrected if the gradient is  
352 well determined (Neuberg and Pointer, 2001). In our case VLP wavelengths are much longer  
353 than the wavelength of the topography and the distortions or corrections would be difficult to  
354 determine.

355

#### 356 **4.4 High-frequency (HF) earthquake location**

357 The 25 HF earthquakes included significant variation in the signal-to-noise ratio from event  
358 to event which negatively affected the ability to pick arrival times and determine the location.  
359 However, detailed examination of manually picked arrival times for *P* phases produced a  
360 subset of 8 of the 25 events which could be located. *S* phases could not be discriminated  
361 (Fig. 4). For these events locations were computed using a uniform half-space model having

362 2.2 km/s (consistent with the model derived in Section 3) using the HYPO71 location  
363 algorithm (Lee et al., 1972). We note that the half space model used here is consistent with  
364 the model used to compute LP and VLP solutions.

365 Locations were initially attempted using a range of alternate models and layered models  
366 which produced substantial variation in hypocentre location. The active source model ( $V_p =$   
367 2.2 km/s) yielded locations beneath the central crater and depth of ~400-500 m. The  
368 modestly higher velocities derived from the Rayleigh wave cross-correlation results yielded  
369 slightly shallower (~100 m) HF earthquake locations. Location results shown in Fig 11 are  
370 summarised in Table 3.

371

## 372 **5. Source inversion**

373 The prior analysis provides well constrained VLP source locations which can now be used to  
374 develop an understanding of the underlying source process. Hence, we proceed next to the  
375 waveform inversion for VLP source excitation. Although many VLP inversions performed  
376 so far (e.g., Ohminato et al., 1998; Nishimura et al., 2000; Hidayat et al., 2002; Chouet et al.,  
377 2003; Kumagai et al., 2003; Chouet et al., 2005; Ohminato, 2006; Aster et al., 2008; Molina  
378 et al., 2008; Waite et al, 2008; Dawson et al., 2011; Chouet and Dawson, 2011; Maeda and  
379 Takeo., 2011; Haney et al., 2012; Maeda et al., 2015a) differ in the inferred source  
380 geometries and interpretation (dykes, sills, connected chambers, dyke-sill composites), they  
381 suggest that the VLP activity represents the elastic response of the rocks in the volcanic  
382 plumbing system to mass transport inside a volcano. While numerous studies deal with VLP  
383 events which are directly related to volcanic explosions (e.g., Chouet et al, 2003; Aster et al.,  
384 2008; Maeda et al., 2015b), there are also reports of VLP activity during periods of eruptive  
385 quiescence (e.g., Saccorotti et al., 2007; Zuccarello et al., 2013).

386 Due to the very long wavelengths (~20–75 km in this study), VLP signals are relatively  
387 insensitive to strong structural heterogeneities present inside volcanoes. This means that  
388 distortions of the wavefield from path effects are minimal, enhancing our ability to reliably  
389 invert for the source mechanisms of these signals. On the contrary, LP signals with the  
390 wavelengths of only several kilometres are severely distorted by both heterogeneity and  
391 pronounced volcano topography (e.g., Neuberg and Pointer, 2000), hence their inversion can  
392 be challenging. Bean et al. (2008) and Trovato et al. (2016) demonstrate through a suite of  
393 synthetic tests that an inaccurate shallow velocity model has a detrimental effect on moment-  
394 tensor solutions of LP sources. Both studies point out that a good waveform fit does not  
395 necessarily imply a correctly recovered source mechanism. In particular, Trovato et al. (2016)  
396 suggest moving towards lower frequencies (if available) and/or using the tilt in the  
397 inversions. In addition to these complications, the signals recorded on White Island, when  
398 filtered in the LP frequency band, are heavily contaminated by noise, largely caused by  
399 oceanic microseisms and wind, which renders them exceptionally challenging for source  
400 inversion. For the outlined reasons, we only perform the source inversion of VLP events.

## 401 **5.1 Method**

402 The seismic ground motion observed at the surface is a convolution of the source (moment  
403 tensor) with the path effect (Green’s functions, GFs). The relationship between the source  
404 components and the surface displacements can be expressed as a linear equation:

405

$$406 \quad \mathbf{u} = \mathbf{G}\mathbf{m}, \quad (6)$$

407

408 where  $\mathbf{u}$  is the data vector containing recorded seismograms, the matrix  $\mathbf{G}$  contains the GFs  
409 for all stations and components and the vector  $\mathbf{m}$  contains the moment tensor components.  
410 The weighted least squares solution of this equation can be written as (Menke, 1984):

411

$$412 \quad \mathbf{m}^{\text{est}} = (\mathbf{G}^T \mathbf{W} \mathbf{G})^{-1} \mathbf{G}^T \mathbf{W} \mathbf{u}, \quad (7)$$

413

414

415 where  $\mathbf{W}$  is a vector of weights for all data traces. The synthetic seismograms corresponding  
416 to this solution are obtained by convolving the solution  $\mathbf{m}^{\text{est}}$  with the GFs (Eq. 6). The  
417 residual  $R$  between the synthetic seismograms and the real data is given by the following  
418 equation:

419

$$420 \quad R = \frac{(\mathbf{u} - \mathbf{G}\mathbf{m})^T \mathbf{W} (\mathbf{u} - \mathbf{G}\mathbf{m})}{\mathbf{u}^T \mathbf{W} \mathbf{u}}. \quad (8)$$

421

422 For the numerical computation of the GFs, the same code (EFISPEC3D) as for the synthetic  
423 test (Section 4.2) was used. The GFs were calculated for a source position obtained with the  
424 semblance method (see Section 4.3) for the 5 best VLP events (highest signal-to-noise ratio;  
425 events 7, 11, 17, 18 and 23 in Table 3), which are clustered within a few hundred m. Like for  
426 the semblance method, here we use a homogeneous model ( $V_P = 2.2$  km/s,  $V_S = 1.272$  km/s  
427 and  $\rho = 2120$  kg/m<sup>3</sup>) for the computations of the Green's functions. As discussed above, the

428 use of a homogeneous model is justified due to the large VLP wavelengths compared to the  
429 dimensions of the experiment.

430 Seismograms on volcanoes are often recorded in the near-field of seismic sources, where tilts  
431 can form a considerable part of the seismic ground motion. Such tilts can have a first-order  
432 effect on horizontal seismograms at long periods (e.g. Rodgers, 1968; Graizer, 2005). Maeda  
433 et al. (2011) proposed a new method of including tilt into seismic inversions by including it  
434 into the GFs. Van Driel et al. (2015) improved and simplified this approach by including the  
435 computation of rotational motion directly in a high-order finite elements method. Here we use  
436 their method to create sets of GFs with and without tilts incorporated in the horizontal  
437 components.

438 In EFISPEC, the rotations are directly obtained from the derivatives of the numerical solution  
439 in the code, rather than approximated with a central differences method at mini-arrays around  
440 the station (as in Van Driel et al., 2015). The gravitational acceleration affecting the  
441 horizontal components are expressed using the spatial derivatives of the ground motion:

$$442 \quad a_i = \frac{1}{2} \left( \frac{\partial u_3}{\partial x_i} - \frac{\partial u_i}{\partial x_3} \right) \cdot g, \quad i = 1,2 \quad (9)$$

443  
444 with the gravitational constant  $g$ , and the displacements  $u_i$ , where  $i = 1,2$  indicate the  
445 horizontal and  $i = 3$  the vertical directions  $x_i$ . The tilt traces are numerically double  
446 integrated and added to the corresponding displacement GFs. As the tilts lead to offsets in the  
447 displacement GFs (as they represent acceleration pulses), the FFT based inversion can  
448 become unstable. We stabilise the result by (i) double differentiating the GFs and (ii)  
449 differentiating the velocity data, the same as in Van Driel et al. (2015).

450 We use 6 different datasets for the inversion: the five single events named above (events 7,  
451 11, 17, 18 and 23) as well as the stack of all 25 VLP events. For the latter, the unfiltered  
452 waveforms of all events were aligned and stacked based on the cross-correlation in the  
453 frequency band 0.03 – 0.125 Hz, with a correlation coefficient threshold of 0.8. This resulted  
454 in a significant noise reduction, particularly on the horizontal components. In preparation for  
455 the inversion, the data is corrected for the instrument response, filtered between 0.04 and  
456 0.125 Hz and integrated to displacement. It is important to mention that the alteration of the  
457 VLP displacement waveforms due to the filtering was very subtle, consistent with our attempt  
458 to minimise the alteration of waveforms prior to inversion. For the inversion, we use all the  
459 seismograms except the faulty east-west components at stations WI02 and WI04 (see above)  
460 and all data is weighted equally.

461

## 462 **5.2 Inversion Results**

463 We first analyse the results of the inversion for the stack. The obtained waveform fits are  
464 plotted in Fig 12, showing generally a good fit between synthetic and real data ( $R = 0.12$ ).  
465 The horizontal components at the stations WIZ, WI14 and WI15 are not well reproduced.  
466 This is likely a consequence of a small error in the source location, which can have a strong  
467 influence on the waveforms at stations close to the source. However, as the large majority of  
468 seismograms fit very well, we consider the solution robust.

469 The resulting full-waveform moment tensor solutions are shown in Fig 13A. The diagonal  
470 moment tensor elements  $M_{xx}$ ,  $M_{yy}$  and  $M_{zz}$  dominate the solution, while the off-diagonal  
471 elements are very small in comparison. The solution shows similar source-time history for all  
472 six moment-tensor components. We use the principal component analysis based on singular  
473 value decomposition (Vasco, 1989) to assess the similarity between the retrieved components

474 and to extract a unique source-time function representative for all the components. The  
475 resulting singular values are shown in Fig 13B, confirming that a single STF is suitable to  
476 describe the inversion results.

477 The resulting source time function (time-history of the source slip) for the stacked traces is  
478 shown in Fig 13C. There is a striking similarity between the slip time-history and recorded  
479 displacement, which confirms that the wavefield is dominated by the near- and intermediate-  
480 field terms (which are directly proportional to the source movement; see Lokmer and Bean,  
481 2010), while the propagation of the far-field  $P$  and  $S$  is negligible in comparison. Table 5  
482 shows the corresponding moment-tensor components for both the stack and the results of the  
483 single VLP events. For the single events, despite the varying amplitude, the relative  
484 contributions of the different moment tensor components are very consistent between events,  
485 supporting a repeating source mechanism. We proceed decomposing the stack solution next.  
486 The inversion including tilts in the Green's functions will be addressed thereafter.

487

### 488 **5.2.1 Moment-tensor decomposition**

489 Following the decomposition method of Vavryčuk (2001), our moment-tensor solution  
490 consists of 75% isotropic, 13% CLVD and 12% double-couple (DC) component. The ratio of  
491 the principal values (PVs) of the stack solution is 1:1.2:1.8, where the eigenvector  
492 corresponding to the largest principal value is quasi-horizontal and pointing approximately to  
493 the North ( $\mathbf{T}$  (E, N, UP) = [0.1 0.98 -0.14]). Moreover, 4 out of 5 individual events yield the  
494 same ratio of principal values, while the remaining event has the ratio 1:1.2:1.9. Such a  
495 robust result makes it possible to infer more detail about the source geometry: specifically,  
496 we search for the best-fit solution assuming a crack with a mixed tensile-shear mode of  
497 failure (Fig 14A), like in Dufumier and Rivera (1997) and Vavryčuk (2001). For  $\alpha = 0^\circ$  the

498 source is a pure DC, for  $\alpha = 90^\circ$  it is a pure tensile crack, while for all other angles it is a  
 499 combination of both. Assuming such a source geometry appears to be plausible because (i)  
 500 the ratio of principal values suggests a quasi-tensile crack and (ii) a crack geometry is the  
 501 most likely candidate for fluid transport inside a shallow volcanic plumbing system (Chouet,  
 502 2003).

503 According to Dufumier and Rivera (1997), the moment-tensor of such a source in the  
 504 coordinate system of principal axis is written as:

$$505 \quad \mathbf{M} = \mu AD \begin{bmatrix} \sin(\alpha)(K+1)-1 & 0 & 0 \\ 0 & K \sin(\alpha) & 0 \\ 0 & 0 & \sin(\alpha)(K+1)+1 \end{bmatrix}, \quad (10)$$

506 where  $A$  is the area of the crack plane,  $D$  is the amount of slip,  $\alpha$  is the inclination of the slip  
 507 from the crack plane, and  $K = \lambda/\mu$  is the ratio of Lamé constants (note that the principal  
 508 values in Eq. (10) appear in ascending order). In order to compare the theoretical principal  
 509 values,  $PV_t$ , from Eq. (10) with our diagonalized moment-tensor solution ( $PV$ ), we normalize  
 510 both sets of principal values with the largest ones and define the L2-misfit ( $R_2$ ) as:

$$511 \quad R_2 = \frac{(\mathbf{PV}_t - \mathbf{PV})(\mathbf{PV}_t - \mathbf{PV})^T}{\mathbf{PV} \cdot \mathbf{PV}^T} \cdot 100\% , \quad (11)$$

512 where

$$513 \quad \mathbf{PV}_t = [PV_t(1)/PV_t(3) \quad PV_t(2)/PV_t(3)] \text{ and}$$

$$514 \quad \mathbf{PV} = [PV(1)/PV(3) \quad PV(2)/PV(3)].$$

515 In order to minimize the  $R_2$  misfit (Eq. (11)), we perform a grid search over the  $(\alpha, K)$ -space,  
 516 where  $\alpha$  ranges from  $0^\circ$  to  $90^\circ$  and  $K$  ranges from 1 to 10.

517 Dufumier and Rivera (1997) and Vavryčuk (2001) point out that the parameter  $K$  is a  
518 property of the fault/crack and it should be inverted for rather than calculated directly from  
519 the velocities of seismic waves in the surrounding medium. The source region is a zone of  
520 weakness and can have significantly different rheological properties than the surrounding  
521 medium. Neglecting this fact may lead to an incorrect determination of the source  
522 parameters.

523 The result of our search for the minimum residual  $R_2$  is shown in Fig 14B. The contours  
524 represent the misfit between the theoretical and observed principal values of the moment  
525 tensor. The absolute minimum is denoted by the blue solid circle and corresponds to  $K = 6.7$   
526 and slip angle  $\alpha = 27^\circ$  (Fig 14A). Since the principal tension axis halves the angle between  
527 the slip direction and the crack normal, it turns out that the crack is dipping at  $67^\circ$  northward  
528 like shown in Fig 15A. However, although our moment-tensor solution was very robust  
529 between 5 events and the stack, the uncertainties in the velocity model, source location and  
530 noise contaminating the data are likely extending possible source candidates to several  
531 percent of the misfit function. Thus, we also show a pure tensile solution, where  $K = 3$  and  $\alpha$   
532  $= 90^\circ$  (Fig 15B). Nevertheless, our solutions show that caution should be exercised when  
533 interpreting moment-tensor solutions for volcanic signals.

534

### 535 **5.2.2 Volumetric change in VLP source**

536 In the previous section, we pointed out the ambiguity in the inferred source mechanism and  
537 the elastic properties of the source region. Here we assess the volumetric change in the VLP  
538 source for both source candidates shown in Fig 15.

539 The isotropic moment ( $\mathbf{M}_0^{\text{ISO}}$ ) of a general moment-tensor ( $\mathbf{M}$ ) is defined as:

540 
$$\mathbf{M}_0^{\text{ISO}} = \frac{1}{3} \text{tr}(\mathbf{M}) \quad (12)$$

541 Combining Eqs. (10) and (12) yields:

542 
$$\mathbf{M}^{\text{ISO}} = AD \sin(\alpha) (\lambda + \frac{2}{3} \mu), \quad (13)$$

543 where  $AD \sin(\alpha)$  is equal to the volume change  $\Delta V$ , which is illustrated in Fig. 14A. This gives  
 544 the relationship between the isotropic moment and the volume change of the source with the  
 545 crack-like geometry:

546 
$$\mathbf{M}^{\text{ISO}} = \Delta V (\lambda + \frac{2}{3} \mu). \quad (14)$$

547 On the other hand, the moment tensor retrieved from the stacked signals (Table 5) is equal to:

548 
$$\mathbf{M} = 1.6 \cdot 10^{12} \begin{bmatrix} 0.69 & 0.03 & -0.02 \\ \mathbf{0.03} & 1 & -0.06 \\ -\mathbf{0.02} & -0.06 & 0.58 \end{bmatrix} \text{Nm}, \quad (15)$$

549 where the isotropic moment (Eq. (12)) is equal to  $\mathbf{M}_0^{\text{ISO}} = 1.2 \cdot 10^{12} \text{Nm}$ .

550 In order to estimate the elastic parameters in the source region and the volume change of VLP  
 551 sources, we use a P-wave velocity ( $V_p = [(\lambda + 2\mu)/\rho]^{1/2}$ ) ranging between 1.3 and 2.3 km/s (see  
 552 Section 3), the density of 2160 kg/m<sup>3</sup> and the retrieved parameter  $K = \lambda/\mu$  to estimate the  
 553 elastic parameters of the source region and the volume change in VLP sources. The results  
 554 for both cases presented in Fig 15 are given in Table 6.

555 The pure tensile and shear-tensile crack mechanisms yield a comparable volume change  
 556 when the same P-wave velocity is assumed. However, the results corresponding to the  
 557 velocities of 2.3 km/s and 1.3 km/s, respectively, give the lower and upper bounds (145 – 450  
 558 m<sup>3</sup>) for the volume change, which encompass the solutions obtained for Stromboli volcano

559 (Chouet et al, 2003), Mt Etna (Zuccarello et al., 2009), Mt Erebus (Aster et al., 2008) and  
560 dike volume change on Popocatepetl volcano (Chouet et al., 2005). However, these authors  
561 dealt with the same ambiguity in the choice of the elastic properties of the medium.

### 562 **5.2.3 Inversion including tilt**

563 For the additional inversion, where tilt was included in the Green's functions, the moment  
564 tensor solution (Table 5) is even more dominated by the isotropic component (86%) and only  
565 small CLVD (4%) and DC (10%) components. Given the noise and uncertainties in the result,  
566 we regard it a nearly isotropic source mechanism; more accurately, only a slightly anisotropic  
567 with the horizontal N-S direction of maximum stress. However, the above decomposition  
568 showed that a wide range of possible candidates is suitable to describe the same moment  
569 tensor.

570 While the fit for the horizontal components at stations WI12, WI14 and WI15 is much better  
571 than for the tilt-free case, the overall waveform fit is not as good as before ( $R = 0.31$ ). The  
572 uncertainties in the source location may provide a reason for this – a relatively small change  
573 in depth can strongly change the amount of tilt in the GFs. Furthermore, large-scale strain can  
574 be converted into local rotations (strain rotation coupling; Van Driel et al., 2012), an effect  
575 not modelled in the GFs. If this coupling is consistent over several stations, it may bias the  
576 obtained results (Van Driel et al., 2015).

577

## 578 **6. Discussion**

579

### 580 **6.1 Semblance and arrival time locations**

581 One goal of this work lies in determining if HF, LP and VLP source processes are from  
582 different spatial locations. In this context, we note that the genetic processes that produce the  
583 three event types may occur in different parts of the hydrothermal/magmatic system. The  
584 location of the three source excitations is also important in the context of our observation that  
585 the VLP source process precedes the LP and HF excitation (see examples in Fig. 3 and 4).  
586 Based on the given locations and uncertainties (Table 3 and Fig 11), we regard the HF, LP  
587 and VLP to have systematically different source positions and wish to examine their  
588 significance in the context of the location procedures and associated uncertainties.

589 We first note that HF source locations appear to be systematically shallower (<1 km) than the  
590 earlier VLP sources (~1 km depth). The VLP sources have highly similar waveforms, which  
591 may be regarded as evidence for a repeating source process and strong rectilinearity that is  
592 expressed in the semblance values.

593 In contrast, the LP sources have very low semblance values which are probably due to a  
594 significant departure in rectilinearity for these events. Indeed, some of the LP events have  
595 unbounded location uncertainties (having error bounds larger than the search grid) (Table 3  
596 and Fig 11 shallow events), suggesting that their locations should be treated with great  
597 caution.

598 In the context of our synthetic results, we regard semblance location uncertainties as due to  
599 several effects including: 1) the elevated noise levels within the LP and VLP signals, 2)  
600 variations in local velocity structure and their departure from the average values found using  
601 active source and ambient noise data, 3) departures from the isotropic source processes, and  
602 4) near-field effect (e.g., see Lokmer and Bean, 2010). The effects of noise are shown  
603 directly in our synthetic analysis (Table 4, compare synthetics case 1 with 3 and case 6 with  
604 8) where the semblance is shown to decrease with increased noise. This effect is also seen by

605 comparison of the signal-to-noise ratios and the computed semblance (see Table 4 and Fig  
606 10). It is clear that the coherence of isotropic components may be significantly contaminated  
607 by noise sources especially for SNR values below 3, so the interpretation of the signals where  
608  $SNR < 3$  should not be attempted.

609 In addition, the semblance values show that some source geometries (e.g., high angle tensile  
610 cracks) may produce significant departures in radial coherence (and lower semblance values).  
611 We suggest here that such cases may be diagnosed if recorded on dense array networks where  
612 discrepancies will tend to reduce the coherence as seen in the semblance computation.

613 The synthetic results suggest, however that azimuthally symmetrical sources (such as a  
614 horizontal tensile crack or a vertical pipe) may generally be indistinguishable from isotropic  
615 sources (compare Table 4 synthetics case 1, 4) from a semblance computation approach. Our  
616 results suggest that a horizontally layered structure may also yield systematic mis-locations in  
617 the data (compare Table 4 synthetic case 1, 7), but note in our case that the velocity structure  
618 is probably better constrained than some other volcanic systems (Section 3). We also note,  
619 based on our synthetic analysis, that vertically oriented dykes will tend to have poorly  
620 retrieved locations, low semblance values and large location errors (Table 4, Case 6 and 8).

621 In this context, we note that VLP sources are generally located at the southern margin of the  
622 Crater and at 1 km depth. This position is roughly coincident with the southern margin  
623 fumaroles (Fumarole 0 and Fumarole 1 in Fig 1), which are thought to tap into gases having  
624 strong magmatic signatures. In addition, the deformation signatures from long term levelling  
625 surveys (Peltier et al., 2009; Fournier and Chardot, 2012) suggest that deformation within the  
626 crater tephra fill is centred at very shallow depth (a few hundred m below the surface), on the  
627 margin of the southern crater wall. Both studies showed that the range of magmatic and  
628 hydrothermal deformation over the past several decades could be explained by the flux of

629 fluids and heat into the shallow hydrothermal system. Interestingly, our HF sources are  
630 located beneath the central crater lake and not the southern margin, which might be viewed as  
631 discrepant with the shallow deformation results. Christenson et al. (submitted), noted that  
632 lake level changes match closely to periods of shallow deformation. Given the modestly  
633 higher velocities for the Rayleigh wave interferometry (2.2 km/s vs 2.4 km/s) yield modestly  
634 shallower locations having the same epicentre cluster, we regard the HF locations as mostly  
635 within the shallow volcano hydrothermal system, possibly caused by repeated failure of  
636 critically stressed rocks in the conduit environment. The results are consistent however with  
637 locations obtained by Nishi et al. (1996) from a short-period array deployment to White  
638 Island. In that study, high frequency earthquakes, referred to as spasmodic and volcano-  
639 tectonic earthquakes, were distributed mostly at very shallow depth beneath the central crater  
640 and present day crater lake. The spectral content for the earlier earthquakes is significantly  
641 higher, peaked at 5-10 Hz, than those observed by our broad-band deployment (peaked at 2-5  
642 Hz) and may have a different genetic provenance. Given the flat response characteristics of  
643 both short-period and broadband seismometers in the common frequency band, we regard the  
644 source process for our HF earthquakes as different from those identified by Nishi et al.  
645 (1996).

## 646 **6.2. Source inversion**

647 We obtained a robust solution for the VLP source inversion, where the most likely source  
648 mechanism is a quasi-vertical, steeply North-dipping tensile crack with a significant amount  
649 of shear movement in the dipping direction. It was pointed out that the elastic constants of the  
650 source region should not be mixed with those from the surrounding medium, as it can result  
651 in retrieving incorrect source parameters. In the case of LP events, we cannot robustly invert  
652 for detail of the source mechanism, so we often *a priori* constrain the solution (e.g., Nakano  
653 and Kumagai, 2005; Lokmer et al., 2007; De Barros et al., 2011), and we suggest that caution

654 be exercised when decomposing the solutions of VLP inversions. When tilt was included in  
655 the computed GFs, the results pointed towards a more isotropic source mechanism.

656 The volume change is estimated to be  $145 \text{ m}^3 - 450 \text{ m}^3$ , which is in agreement with the  
657 results reported at different volcanoes (see references in the last paragraph of Section 5.2.2).

658 In all these examples, the VLP pulses are attributed to the ascent of a gas slug. Interestingly,  
659 Zuccarello et al. (2009) estimated that the cumulative volume change of VLP sources  
660 contributed to only about 5% of the total gas emission on Mt. Etna volcano, a result that  
661 points to uncertainties in the mass balance between gas transmission at depth and subsequent  
662 discharge at the surface. However, the reason might be more trivial: if there are significantly  
663 lower frequencies present in the recorded VLP signals, we would not be able to recover them  
664 truthfully due to the band-limited nature of our instruments. If this is the case, then the  
665 retrieved source-time function (Fig 13C) does not represent the true time-history of the slip in  
666 the source, but it is rather its filtered representation, and may have significantly lower  
667 amplitude than the true process. This scenario is very likely for slow processes of small  
668 magnitude, whose duration is comparable to or longer than the lower frequency limit of the  
669 recording instrument. Maeda et al (2015a, b) recognized this problem and performed a search  
670 for the parameters of elementary functions (ramp, Gaussian pulse and triangle, see Fig 7 in  
671 Maeda et al., 2015a) whose filtered version corresponded well to the retrieved source-time  
672 function. Here we suggest that such a possibility should not be ignored and caution should be  
673 exercised when directly linking the retrieved source function to a source process.

674

### 675 **6.3 Interpretation of White Island fluid flux from coupled earthquakes**

676 We next examine the process that might generate the three coupled event types based on their  
677 timing and location. We noted earlier that the main VLP event onset precedes the HF and LP

678 sources by a few seconds. The HF and LP sources arrival times are difficult to distinguish in  
679 our case, but we do note that the LP source extends in time beyond the coda for the HF event  
680 type. In addition, the VLP is deeper (~1 km) than both the HF and LP sources. We do not  
681 have strong confidence in the LP source locations (Fig 11), due to the strong departures from  
682 radial particle motions (see synthetics section 4.2), but propose that the HF source process is  
683 associated with very shallow portion of the crater system.

684 We place the event types in the context of an existing conceptual model for the White Island  
685 volcanic system (Fig 16), composed of a shallow magma storage region overlain by a shallow  
686 convecting hydrothermal system (e.g., Christenson and Wood, 1993; Werner et al, 2008).  
687 The hydrothermal system likely includes strong permeability contrasts from inherent rock  
688 porosity and the fragmental vent system, which is cut by fumarolic pathways to the surface  
689 and into the crater lake. The hydrothermal system includes a single phase vapour within the  
690 high flux pathways and near the magma interface. Away from the fumarolic pathways and at  
691 shallower depth, the single phase gas is replaced by a two phase system including fluids and  
692 gas, which in turn gives way to a single phase fluid at the crater lake near surface water table.  
693 Convection and persistent degassing is driven by heat and gas from the shallow magma  
694 system (Fig 16).

695 We propose the VLP source process as possibly caused by the release of gas and fluids that  
696 have ponded behind the top of the magma conduit (Fig 16). The ponding may be akin to a  
697 bubble trap (e.g., Adelstein et al, 2015) owing to a mostly impermeable magma carapace (Fig  
698 16), which has been inferred at White Island (Cole et al 1999) based on geophysical  
699 constraints from eruptive activity in 1977 (Clark and Otway, 1989; Houghton and Nairn,  
700 1989). Release of slowly pressurized fluids from the bubble trap may cause upward  
701 migration and very long period collapse and recovery behind the trap. It is uncertain if this  
702 pressurisation step is linked to the observed 30 second onset of the VLP signal, shown in the

703 low noise examples in Fig 4, which in our case may be the slow deformation onset seen in  
704 their Fig 2 (Maeda et al., 2015b). In this case, the primary VLP oscillation (rupture onset)  
705 may be observed as a strongly isotropic radiation pattern, consistent with rupture of the  
706 carapace. The rupture would include both tensile and shear crack components (Fig 15) as the  
707 carapace fails. Hot ductile conditions near the magma would promote rapid healing of the  
708 carapace after the VLP and the introduction of new gas would quickly re-pressurise the top of  
709 the magma system. The time frame of rupture and healing in this case may be on the order of  
710 the inter-event times for the coupled earthquakes (about 2 to 3 hours). The moment tensor  
711 inversion and volume computation are computed assuming elastic conditions in the source  
712 medium. It is uncertain if such conditions would be satisfied for the seismological time  
713 frames seen for VLP at White Island. At Kilauea volcano, Hawaii, zones of hot weak rock  
714 (Dawson et al., 1999) are interpreted from strong  $V_p/V_s$  contrasts (and attendant high Poisson  
715 ratios) found from travel-time tomography. Ohminato et al, (1998) suggests that high  
716 Poisson ratios ( $\sim 0.33$ ) may be required to explain moment tensor solutions for tensile crack  
717 excitations inferred at that volcano. At White Island, the inferred magma carapace would  
718 have similar rock mechanical conditions and could be the locus of VLP source processes.  
719 We note however, that our moment tensor inversion (Section 5.2.2) does not provide  
720 information about the elastic rock conditions at the source. Instead, these may be viewed as  
721 inversion parameters for discontinuities (e.g., faults). This is highlighted in Dufumier and  
722 Rivera (1997) and Vavrycuk (2001). In fact, Vavrycuk (2001) used the Poisson ratio as a  
723 parameter to separate realistic and unrealistic source mechanism solutions. If we accept that  
724 the VLP represents a fault-like rupture of the magmatic carapace, then the rupture process  
725 may be elastic for time frames similar to the VLP and ductile for longer timeframes (Melosh,  
726 1980).

727 Meanwhile, as the gas slug proceeds to shallower depth, the onset of LP and HF excitations  
728 begin. We emphasise that the short time delay between the primary VLP and LP/HF source  
729 onsets may preclude gas slug migration as a viable mechanism for the shallow seismicity.  
730 Gas slug migrations times (James et al., 2006) for gas through fluids are probably an order of  
731 magnitude too slow to account for the excitation of source positions as much as 500 m apart.  
732 If instead the gas pulse is propagating through a single phase gas (e.g. Christenson et al,  
733 2010), then the propagation times might be faster. For the former, the fluid/gas pressure  
734 above the ascending gas slug may act to perturb critically stressed portions of the shallow  
735 hydrothermal system leading to HF (Nishi et al, 1996) and LP activity (e.g., Chouet, 1993;  
736 Neuberg et al, 2000) in the shallow hydrothermal system. For the latter, the rapid propagation  
737 of the gas slug may directly induce the HF seismicity as discussed below. We note that  
738 significant overpressures/underpressures may be achieved from the passage of gas slugs  
739 through conduit flares (James et al, 2006). In our case, LP and HF activity may not be a  
740 direct result of the gas slug propagating through a flared conduit. Instead, we surmise that the  
741 locus of shallow seismicity may be focused at points of stress concentration and may include  
742 flares or constrictions within the hydrothermal conduit system. Similarly, the HF and LP  
743 events are not interpreted as an example of the hybrid event type first described by Lahr et al.  
744 (1994) due to the variable timing between HF and LP sources . This may be illustrated by  
745 stacking coherent LP waveforms for a given event (example completed using event 11 in Fig  
746 3 and Table 1) and then noting small migrations (a few tenths of a second) in the onset of the  
747 HF component (Fig 17). The LP-HF timing offset might occur either by the variable  
748 propagation speeds of HF body waves and LP surface waves to the dense array of stations, or  
749 may alternatively imply a different physical location for the two wave types. The small  
750 timing offsets suggest the later in our case. In addition, we note that the LP source process,  
751 while coherent station-to-station (Fig 17A) are not strongly coherent from event-to-event (Fig

752 17C), based on an analysis of the waveform cross correlation. The lack of coherence can be  
753 partially explained by the strong effects of noise on the LP source (compare Fig 17C to Fig  
754 3), but significant variation in waveform coherence is seen for periods having low noise  
755 levels (see variation in coherence for events 15-23 in Fig 17 C). By contrast, coherence for  
756 VLP events is very high (Fig 17D) for all events of the sequence. This implies that the VLP  
757 events are repetitious in nature while the LP component represents a more variable source  
758 process. Results suggest that the LP source process seen here may be different from those  
759 observed by Sherburn et al., (1998), based on the strong event similarity seen in the earlier  
760 study.

761 We also consider the coupled event types in the context of the possible modes of material  
762 failure at White Island. For example, Heap et al, (2015) noted that hydrothermally altered  
763 lavas at White Island will favour dilatant failure to depths of 2 km while ash tuffs convert  
764 from dilatational to compactant modes at depths of ~250 m. Further, Heap et al. (2015)  
765 suggest that White Island ash tuffs, which are prone to low pressure failure from pore  
766 collapse, may promote small, long rupture duration events such as LP seismicity (Bean et al.,  
767 2014) at very shallow depths. However, it is difficult to envisage the repeating failure of  
768 coupled LP and HF events for two separate rock types distributed within a few seconds of  
769 each other and this may be implausible for these earthquakes. We regard the mechanical  
770 failure modes outlined by Heap et al. (2015) as more appropriate for VT and LP type activity  
771 found by Nishi et al., (1996) and Sherburn et al. (1998), which may have different genetic  
772 source process than for the coupled events described here. In our case, we favour failure of  
773 the magma carapace as the source for VLP, resonance associated with fluid migration as the  
774 source for shallow LP activity and the rapid failure of critically stressed conduit materials as  
775 the source for HF events. We note similarities between the observations in our study, and  
776 those at Satsuma-Iwojima volcano in Japan (Ohminato, 2006) which documented VLP

777 coupled with LP burst activity. We note however that Ohminato (2006) suggested a very  
778 difference source process, linking the activity to the repeated excitation of boiling water to  
779 the very shallow hydrothermal system. This contrast in interpretation illustrates the need for  
780 detailed understanding of source location and process in the interpretation of seismic signals  
781 from volcanic systems.

782

## 783 **7. Conclusions**

784

785 We completed a detailed analysis of 25 coupled HF, LP and VLP earthquakes recorded on a  
786 dense array deployment at White Island between 19-21 August 2011. We find that the VLP  
787 component precedes the HF and LP components of the activity by a few seconds. The VLP  
788 signals have generally deeper locations (~1.0 km) than the HF and LP events (<1 km), based  
789 on locations using waveform semblance and arrival time earthquake locations.

790 The source processes producing the LP, VLP and HF source components are likely to be  
791 genetically linked to the rapid flux of fluids (gas and/or liquid phases) from the magma  
792 system into the shallow hydrothermal system. Moment tensor inversion of the VLP events,  
793 which are centred below two high gas discharge fumaroles, is consistent with a high angle  
794 north dipping and east-west striking tensile crack. Each tensile excitation includes a modest  
795 shear component directed down the dipping plane.

796 Each VLP event is interpreted as a repeating excitation of this stationary crack above a  
797 shallow magma storage region. We surmise that, at low excess pressures, the crack releases  
798 gasses passively, but for strong gas flux periods, the gas release produced discrete VLP-HF  
799 and LP source excitations. The volume change associated with the VLP component is about

800 145-450 m<sup>3</sup> while the moment release is  $\sim 1.2 \times 10^{12}$  Nm. Each event may therefore reflect the  
801 release of gas from the magmatic hot ductile carapace into the overlying hydrothermal  
802 system. The HF and LP components may represent the seismic response of fluids  
803 pressurisation in the overlying hydrothermal system. The LP source may either be associated  
804 with resonant fluid migration (e.g., Chouet, 1996; Neuberg et al, 2002), or less likely, the  
805 failure of low tensile strength ash tuff materials (Heap et al., 2015; Bean et al., 2014). The  
806 HF earthquakes may result from either brittle failure of higher strength lava flow materials  
807 (Heap et al., 2015), or more likely the rapid failure of critically stressed rock in the  
808 hydrothermal conduit. Gas slug ascent rates within a fluid filled conduit are probably too  
809 slow (James et al., 2006) to produce LP/HF seismicity at shallow depth in our case but slug  
810 ascent within a pre-existing single phase gas column may promote more rapid gas ascent. In  
811 the former case, the shallower seismicity is possibly due to either stress transmission directly  
812 through the rock or via fluid pressure transfer in the hydrothermal conduit system. In the  
813 latter case, gas ascent itself may initiate HF and LP activity. Future work will examine  
814 additional occurrences of VLP activity specifically associated with eruptive activity in  
815 October 2013.

816

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828

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1134 Table 1: VLP filtered at 0.04-0.125 Hz. N is the event number and locations with error limits (in meters given  
1135 in New Zealand Map Grid coordinates). Lateral and depth errors ( $er_l(z)$ ) are in meters based on 2% error limit  
1136 from the maximum semblance. SNR is the computed signal to noise ratio.

1137 N	event easting	northing	depth	$er_l(z)$	semblance	SNR
1138 1	2879980+/-170	6400000+/-130	880	660-1040	0.4774	2.3859
1139 2	2879820+/-140	6400060+/-140	840	640-1000	0.5497	2.3355
1140 3	2880020+/-150	6399840+/-130	940	740-1160	0.4909	1.9413
1141 4	2880320+/-020	6400360+/-060	740	560-940	0.4730	2.7645
1142 5	2880420+/-040	6400140+/-100	860	640-1060	0.4821	2.4670
1143 6	2880100+/-130	6399880+/-140	920	720-1120	0.5278	2.3567
1144 7	2880220+/-090	6400080+/-100	840	620-1000	0.6310	3.6471
1145 8	2880000+/-110	6400100+/-130	900	700-1100	0.5720	2.8050
1146 9	2880200+/-130	6400000+/-150	920	700-1100	0.5983	2.3324
1147 10	2880360+/-030	6400100+/-090	820	640-1020	0.5263	2.9952
1148 11	2880360+/-010	6400160+/-090	960	760-1180	0.5309	4.3236
1149 12	2880060+/-160	6399800+/-120	920	720-1120	0.5180	2.8768
1150 13	2880380+/-090	6400040+/-070	680	540-860	0.3955	1.5893
1151 14	2880260+/-110	6400060+/-090	920	700-1100	0.6862	2.8702
1152 15	2880360+/-020	6400080+/-090	960	720-1160	0.5961	3.4776
1153 16	2879760+/-200	6399740+/-170	1120	880-1120	0.5654	2.4453
1154 17	2880060+/-140	6399900+/-160	1040	820-1280	0.6724	4.5016
1155 18	2880000+/-160	6399980+/-150	980	760-1180	0.6420	4.5097
1156 19	2879800+/-140	6400020+/-130	1180	760-1180	0.4870	2.8766
1157 20	2879980+/-190	6399820+/-150	1180	940-1440	0.5718	1.3369
1158 21	2880000+/-180	6399940+/-140	960	740-1160	0.6191	3.2514

1159	22	2880280+/-050	6400080+/-080	920	700-1100	0.6462	3.5423
1160	23	2880240+/-070	6400080+/-090	960	740-1160	0.7133	4.3724
1161	24	2880300+/-070	6400080+/-090	840	660-1040	0.5877	3.3617
1162	25	2880020+/-140	6399920+/-090	520	400-640	0.1803	1.2227
1163							
1164							
1165							
1166							
1167							
1168							
1169							

1170 Table 2: LP phase filtered 0.5-1.1 Hz. Error limits 50% of maximum semblance. See table 1.

1171	N	event easting	northing	depth	er_l(z)	semblance S/NR	
1172	1	2880160+/-240	6400400+/-190	280	20-920	0.1711	2.4994
1173	2	2880200+/-190	6400440+/-110	300	60-620	0.0792	2.1020
1174	3	2880300+/-110	6400500+/-070	320	160-460	0.0374	1.7694
1175	4	2880220+/-290	6400480+/-160	400	60-1120	0.1751	2.2955
1176	5	2880100+/----	6400300+/----	000	-----	-0.1121	1.2562
1177	6	2880040+/----	6400220+/----	060	-----	-0.1245	1.3046
1178	7	2880220+/-280	6400480+/-140	400	80-1180	0.1221	2.8868
1179	8	2880180+/-240	6400444+/-140	260	140-240	0.1135	1.9039
1180	9	2880320+/-290	6400440+/-190	440	60-1080	0.1631	2.7176
1181	10	2880200+/-350	6400460+/-190	420	40-1400	0.2220	2.5738
1182	11	2880340+/-190	6400400+/-120	300	100-600	0.0697	2.8215
1183	12	2880260+/-360	6400480+/-200	560	100-1760	0.2255	4.0386
1184	13	2880180+/-210	6400420+/-100	280	40-720	0.1036	2.0479
1185	14	2880220+/-160	6400440+/-090	280	80-560	0.0689	2.4666
1186	15	2880320+/-330	6400440+/-210	440	80-1340	0.2063	3.8058
1187	16	2880160+/-310	6400420+/-180	420	20-1560	0.2535	6.5132
1188	17	2880320+/-430	6400580+/-240	960	220-1980	0.1908	7.8226
1189	18	2880280+/-310	6400480+/-170	500	100-1280	0.1570	4.2392
1190	19	2880220+/-250	6400440+/-130	320	40-800	0.1289	2.5791
1191	20	2880340+/-380	6400520+/-250	600	120-1840	0.2360	5.5220
1192	21	2880180+/-290	6400420+/-240	280	20-1040	0.2150	2.9094
1193	22	2880360+/-260	6400460+/-180	420	120-1260	0.1299	2.8239
1194	23	2880180+/-270	6400420+/-190	280	00-780	0.1536	2.1955
1195	24	2880120+/-----	6400400+/-----	180	-----	-0.0600	1.5858
1196	25	2880000+/-----	6400140+/-----	20	-----	-0.1239	1.0630

1197

1198

1199 Table 3: HF earthquake locations from hypo71 (Lee et al, 1972). Locations as described in Table 1, with lateral  
 1200 standard errors given by (seh) and depth standard errors given by (sez). The average root-mean-square (rms)  
 1201 has units of seconds.

1202		easting	northing	depth	seh(m)	sez(m)	rms
1203	7	2880002	6400232	150	141	600	0.10
1204	11	2880082	6400417	430	071	300	0.06
1205	12	2880146	6400081	1350	212	1200	0.07
1206	13	2880042	6400141	340	071	500	0.07
1207	14	2880117	6400416	540	141	300	0.07
1208	15	2880064	6400218	500	071	400	0.07
1209	16	2880072	6400395	260	071	200	0.05
1210	17	2880068	6400307	460	071	300	0.06

1211

1212

1213 Table 4: Test of the waveform semblance location using a range of synthetic seismograms test cases. Eight  
 1214 cases examine the performance of the semblance method having the effects of variable source models, noise  
 1215 levels and noise levels. SynthN is the case number including the location (easting, northing and depth and  
 1216 description of the case. Synthetic seismograms are produced using the methods outlined in Section 4.2. The  
 1217 location recovered from the semblance analysis is computed using the methods outlined in Section 4.1.

1218 Case

1219	synthN	event easting	northing	depth	description
1220	sembN	event easting	northing	depth	semblance
1221					
1222	Homogeneous model				
1223	synth1	2880132	6400282	1000	explosive source
1224	semb1	2880200	6400300	800	0.9733
1225	synth2	2880132	6400282	2000	explosive source
1226	semb2	2880200	6400200	1900	0.9568
1227	synth3	2880132	6400282	1000	explosive source with noise
1228	semb3	2880100	6400200	900	0.8473
1229	synth4	2880132	6400282	1000	horiz. tensile crack

1230 semb4 2880200 6400300 800 0.9762  
1231 Layered model  
1232 synth5 2880132 6400282 1000 explosive source  
1233 semb5 2880200 6400300 600 0.9724  
1234 synth6 2880132 6400282 1000 oblique tensile crack#  
1235 semb6 2880600 6399500 2100 0.2652  
1236 synth7 2880132 6400282 1000 horiz. tensile crack  
1237 semb7 2880200 6400300 600 0.9749  
1238 synth8 2880132 6400282 1000 oblique tensile crack# w noise  
1239 semb8 2880600 6399600 900 -0.0032  
1240 # oblique tensile crack oriented at strike 25° and dipping 75°.  
1241

1242 Table 5: Results of the moment tensor inversion for the stack and 5 single events. The value  
1243 A corresponds to the maximum moment tensor component, the other values denote the  
1244 relative amplitudes of the individual moment tensor components  $M_{ij}$ . The last line shows the  
1245 result for the inversion including tilt.

1246

VLP Event	A (Nm)	$M_{xx}$	$M_{yy}$	$M_{zz}$	$M_{xy}$	$M_{xz}$	$M_{yz}$
<b>Stack</b>	$1.60 \times 10^{12}$	0.69	1.00	0.58	0.03	-0.02	-0.06
<b>7</b>	$2.13 \times 10^{12}$	0.63	1.00	0.55	0.05	-0.02	-0.06
<b>11</b>	$1.96 \times 10^{12}$	0.67	1.00	0.56	0.03	-0.02	-0.06
<b>17</b>	$2.11 \times 10^{12}$	0.66	1.00	0.56	0.04	-0.02	-0.06
<b>18</b>	$1.83 \times 10^{12}$	0.68	1.00	0.57	0.04	-0.02	-0.06
<b>23</b>	$1.98 \times 10^{12}$	0.67	1.00	0.56	0.04	-0.02	-0.06
<b>Stack*</b>	$5.98 \times 10^{11}$	0.95	1.00	0.73	0.06	-0.03	0.02

1247

1248

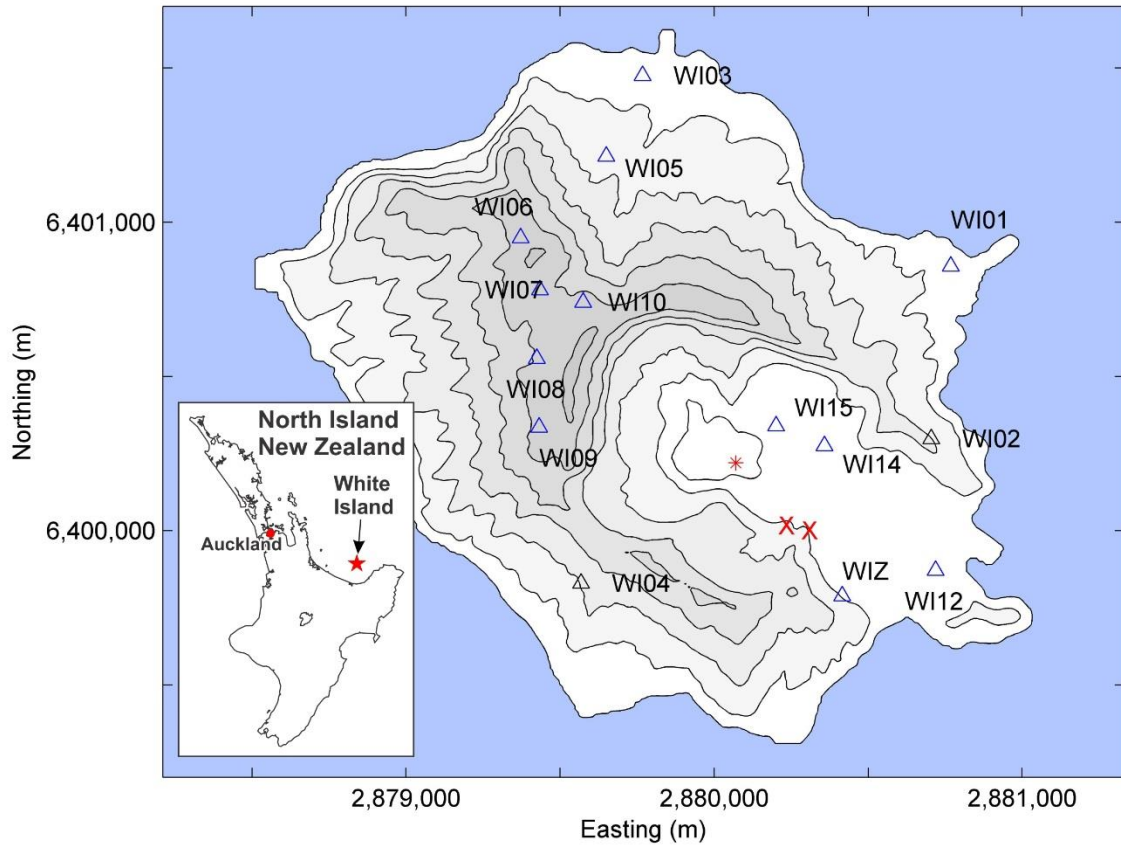
1249 Table 6: Parameters obtained for the VLP stack solution for a mixed shear-tensile crack  
1250 mechanism as described by Dufumier & Rivera (1997) and Vavrycuk (2001). To calculate  
1251 the volume change, the isotropic moment  $M_0^{ISO} = 1.2 \times 10^{12}$  Nm from the moment tensor  
1252 inversion was used. To estimate elastic parameters and volume change of the source, we  
1253 retrieved the parameter  $K$  by varying  $V_p$  at an assumed density of  $2160 \text{ kg/m}^3$  (see Fig 15).

$v_p$ [km/s]	$\rho$ [kg/m <sup>3</sup> ]	$K$	$\lambda$ [GPa]	$\mu$ [GPa]	$M_0^{ISO}$ [Nm]	$\Delta V$ [m <sup>3</sup> ]
1.3	2160	6.7	2.8	0.4	$1.2 \cdot 10^{12}$	391
2.3	2160	6.7	8.7	1.3	$1.2 \cdot 10^{12}$	125
1.3	2160	3	2.2	0.7	$1.2 \cdot 10^{12}$	450
2.3	2160	3	6.9	2.3	$1.2 \cdot 10^{12}$	142

1254

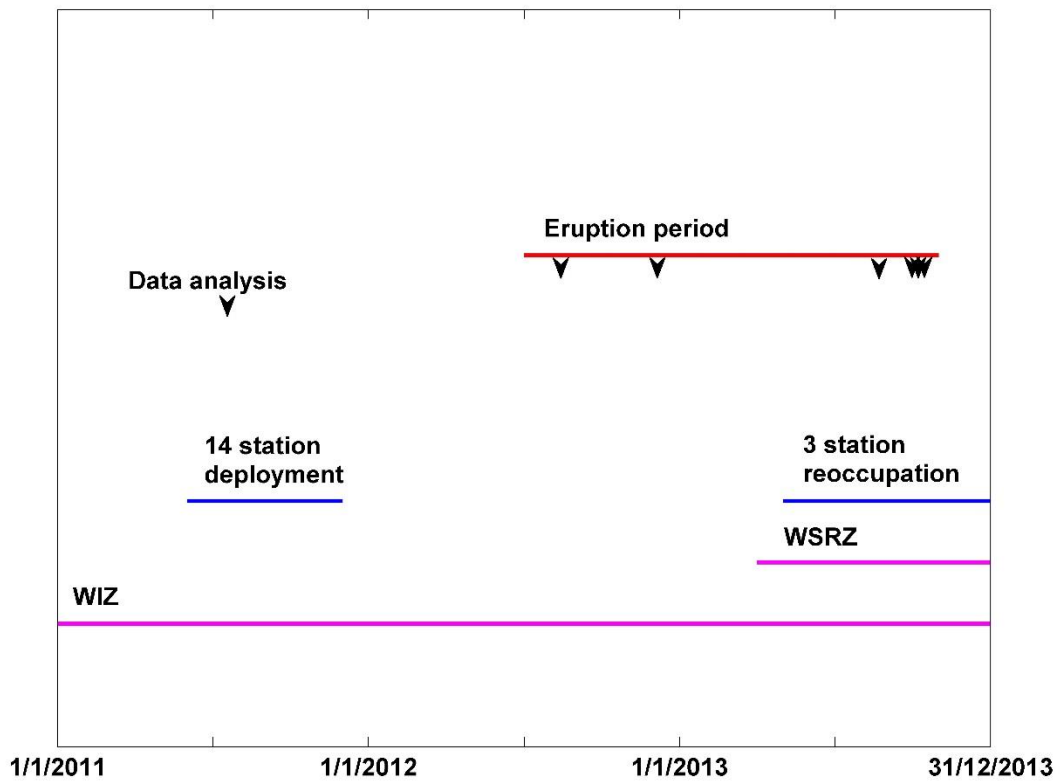
1255

1256



1258

1259 Fig 1: Map showing White Island with location of permanent seismic station (WIZ) and  
 1260 temporary portable seismic stations (open blue triangles) deployed in 2011 Stations WI02 and  
 1261 WI04 (black triangles) had issues with horizontal component sensors and hence were not  
 1262 used for the semblance computation. The site of small scale eruptive activity in 2012-2013 is  
 1263 marked by stars while the red X's mark the location of high discharge fumaroles. Contour  
 1264 interval is 50 m.

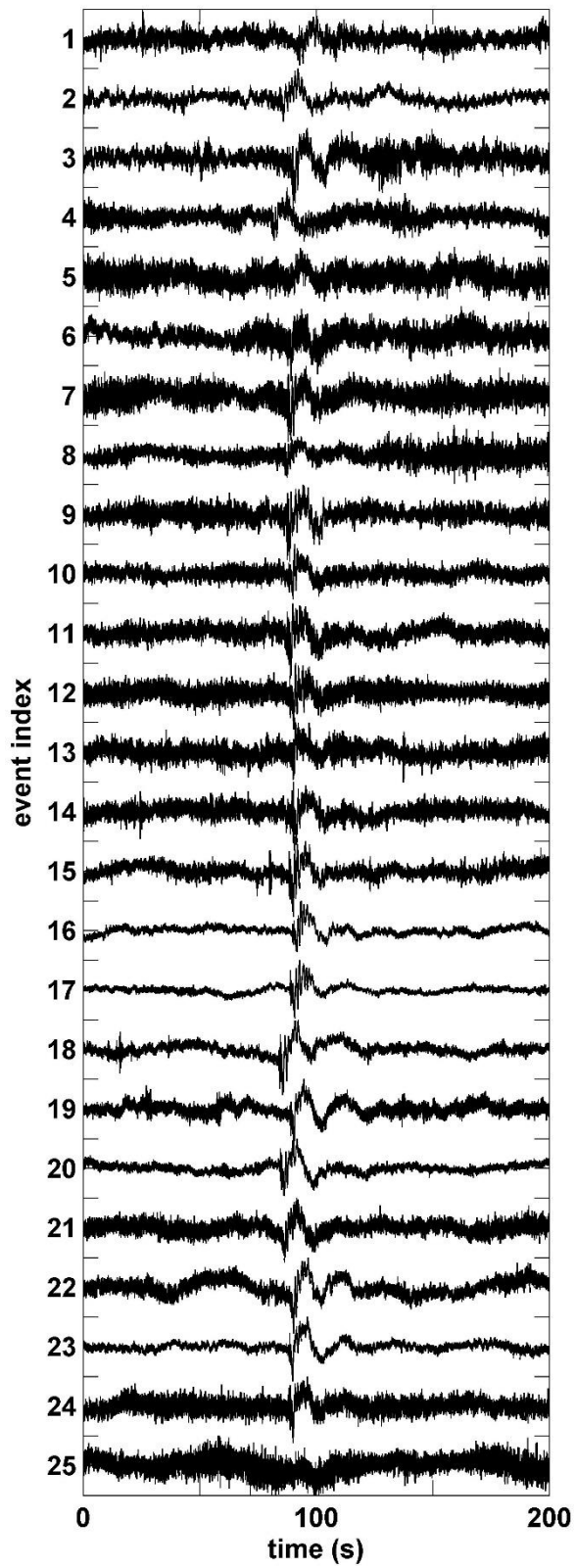


1265

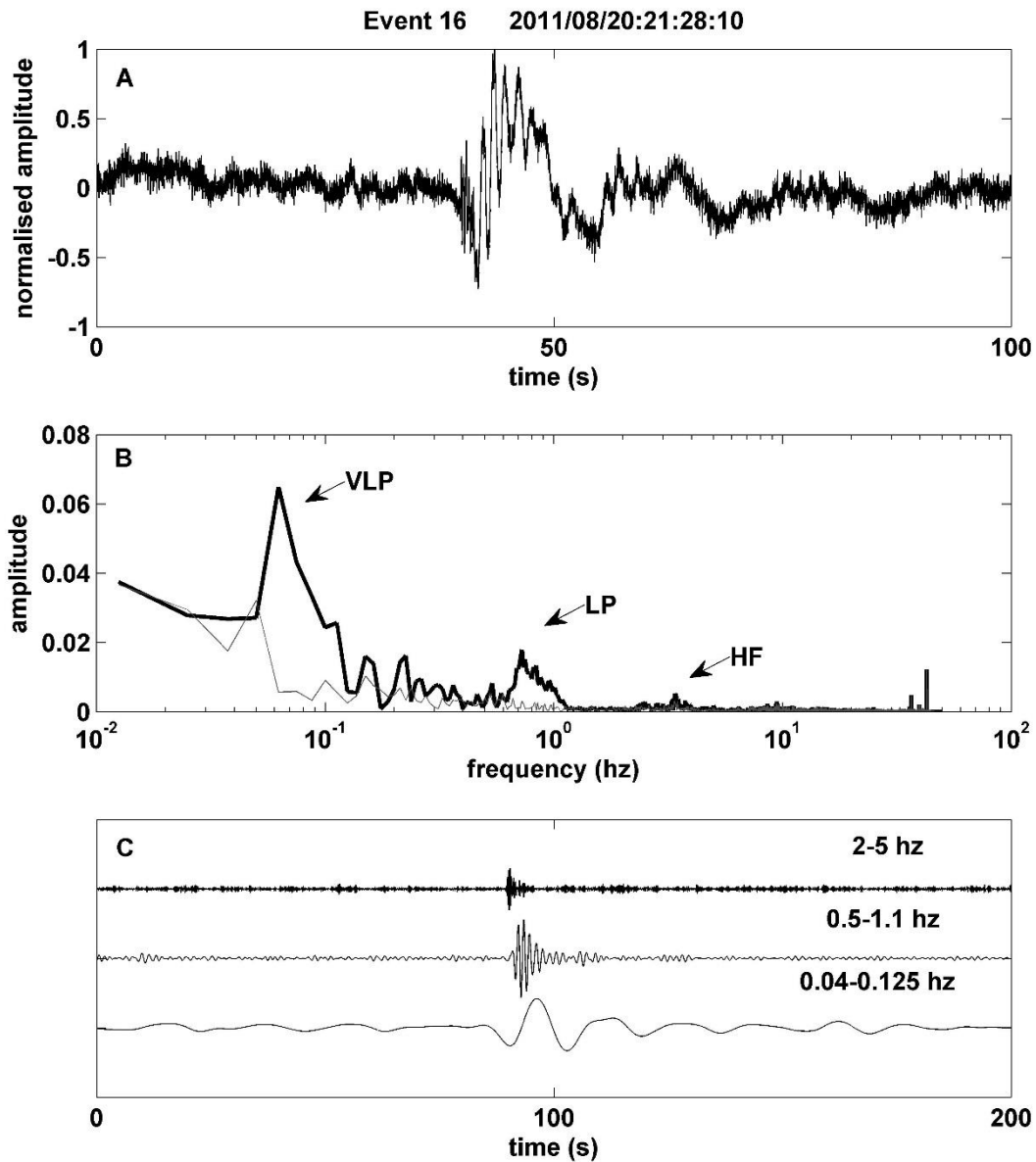
1266 Fig 2: Timeline of volcanic unrest (red line) and individual eruptive events (black arrows).

1267 The analysis presented here is shown by the black arrow including the period of the 14 station

1268 deployment in blue. The permanent station operation is shown in magenta.

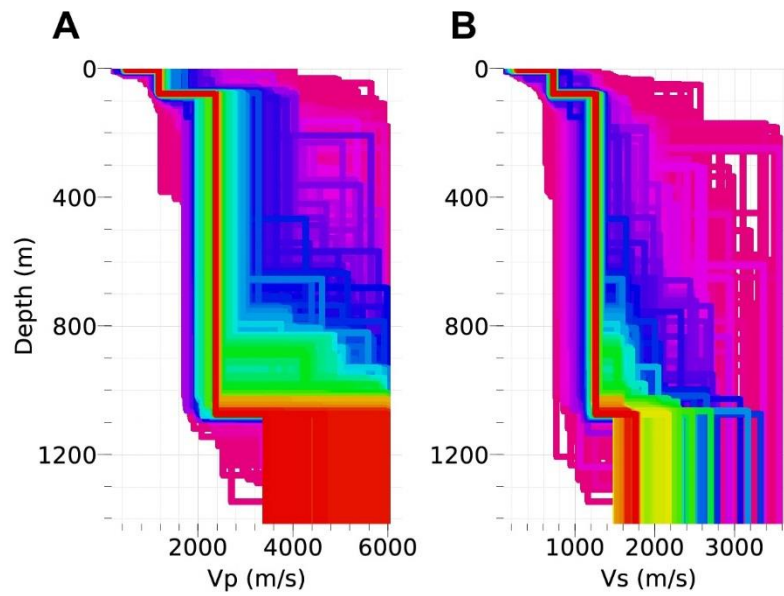


1270 Fig 3: Unfiltered normalised velocity waveforms for an earthquake swarm occurring 19-21  
1271 August 2011 on station WIZ. The events are identified based on their VLP signal but also  
1272 include variable long period and high frequency components.



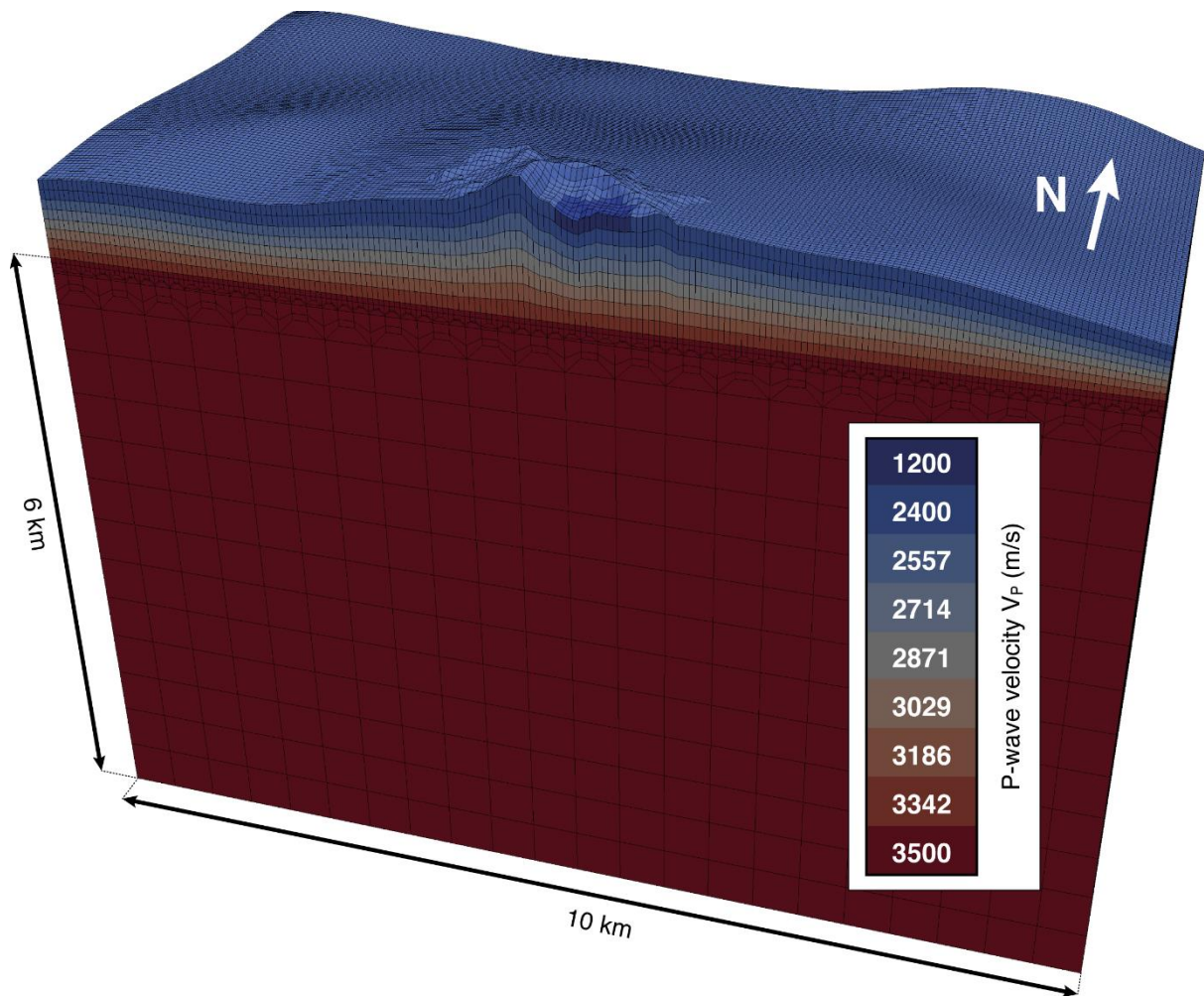
1273  
1274 Fig 4: Example waveform and spectra for event 16 showing multi-phase character of the  
1275 sequence. Panel A shows the raw normalised seismogram, panel B shows the associated  
1276 spectrum with VLP (0.04-0.125 Hz), LP (0.5-1.1 Hz) and HF (2-5 Hz) components. The

1277 spikes observed at >40 Hz are probably associated with wind-noise. The pre-event noise is  
1278 shown as the grey line. Zero-phase filtered seismograms in the selected pass bands are  
1279 shown in Panel C.



1280

1281 Fig 5: Example inversion of Rayleigh dispersion from the White Island dense array  
1282 deployment for  $V_p$  (A),  $V_s$  (B). Warm colours represent models with lower misfits..

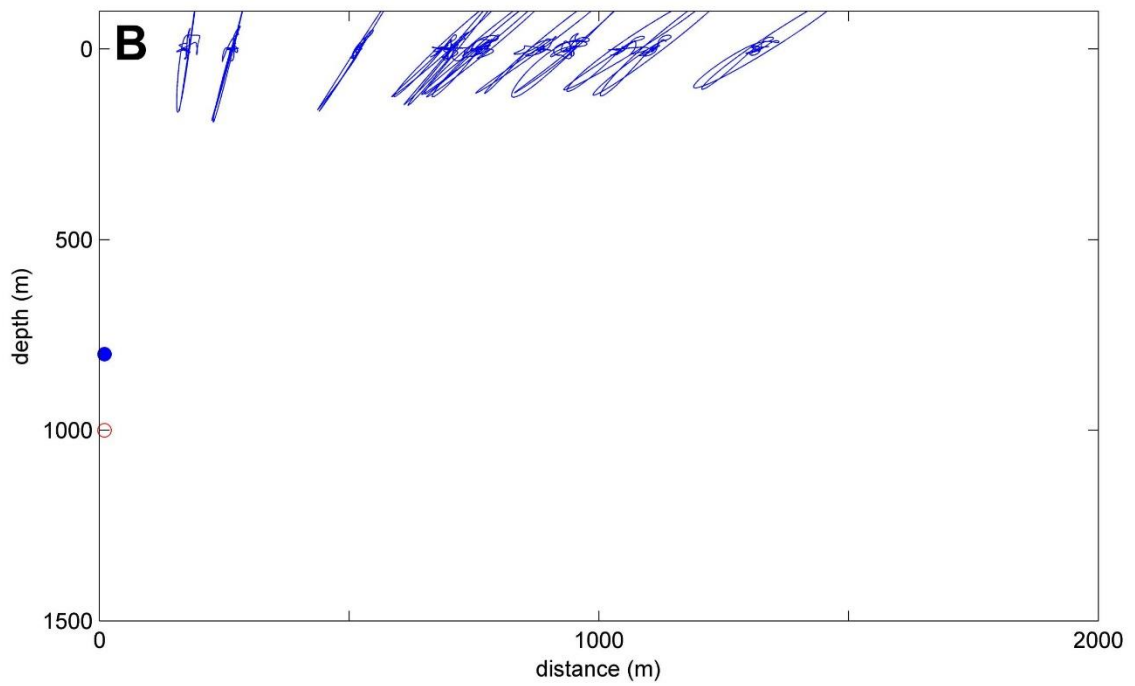
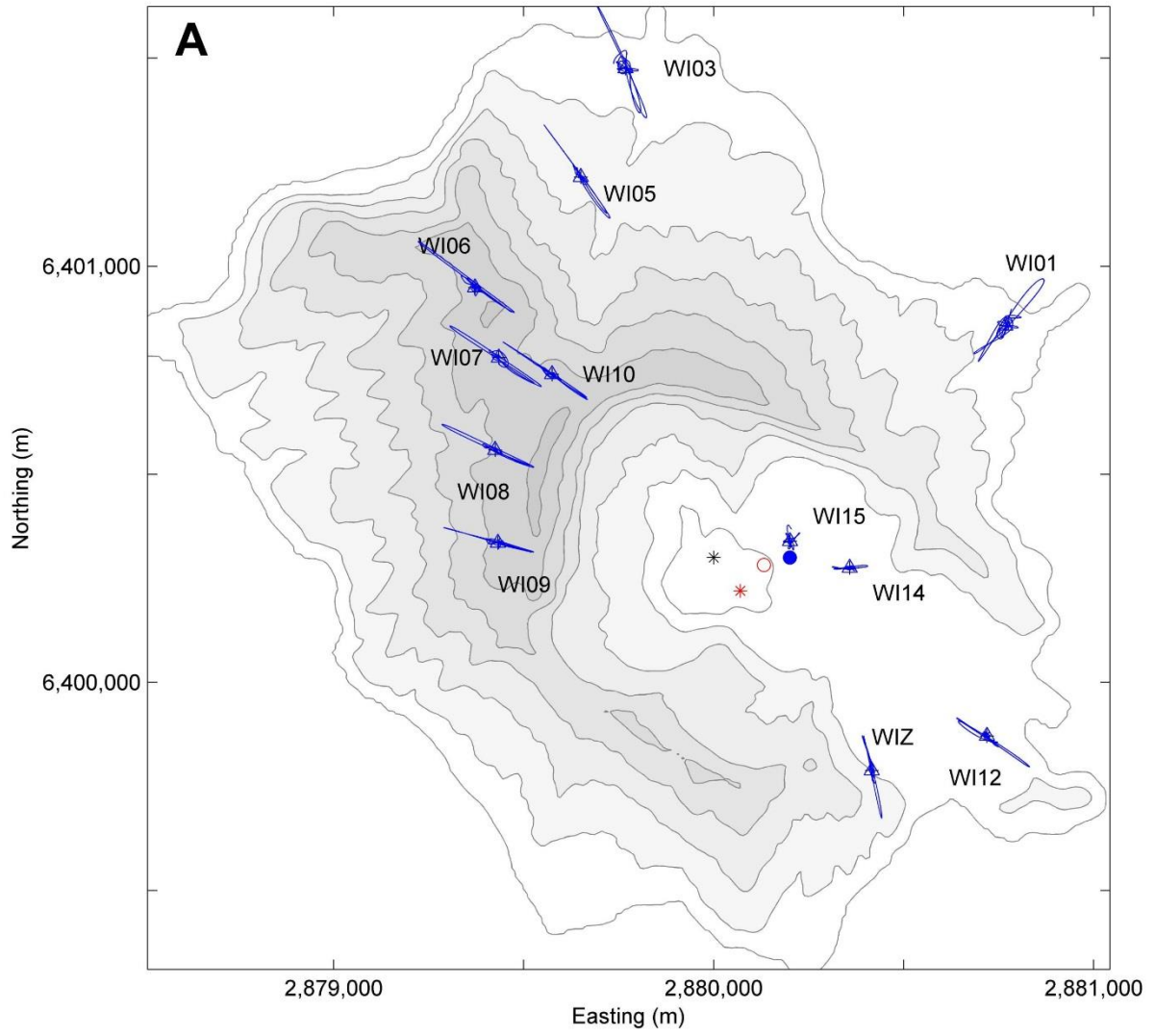


1283

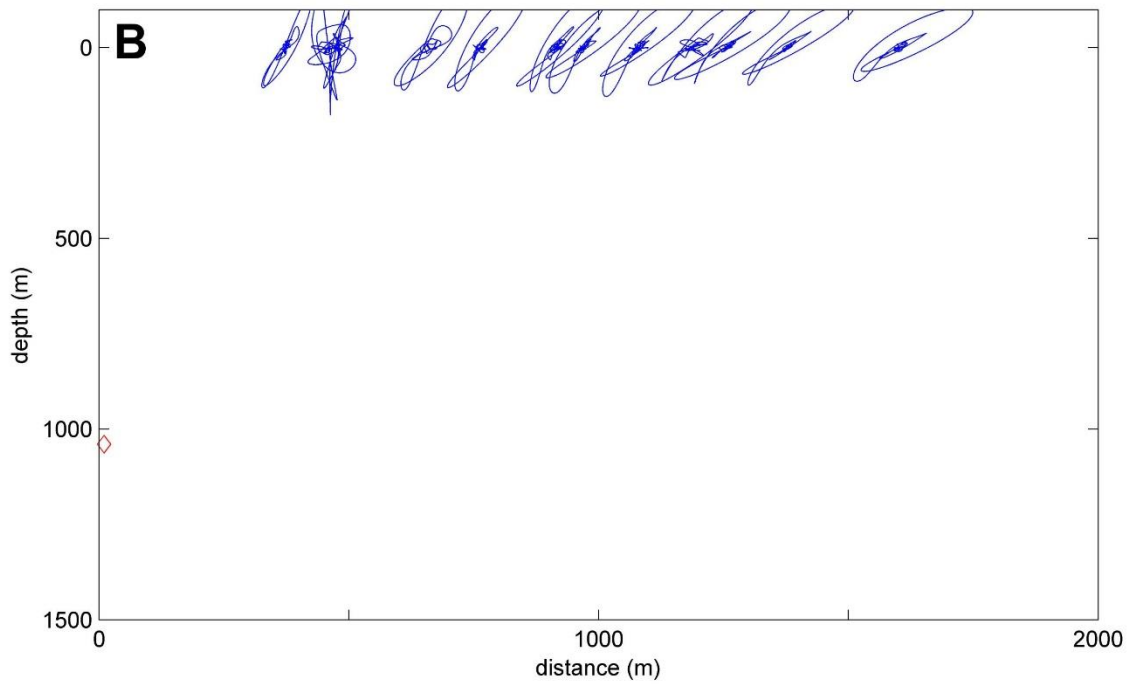
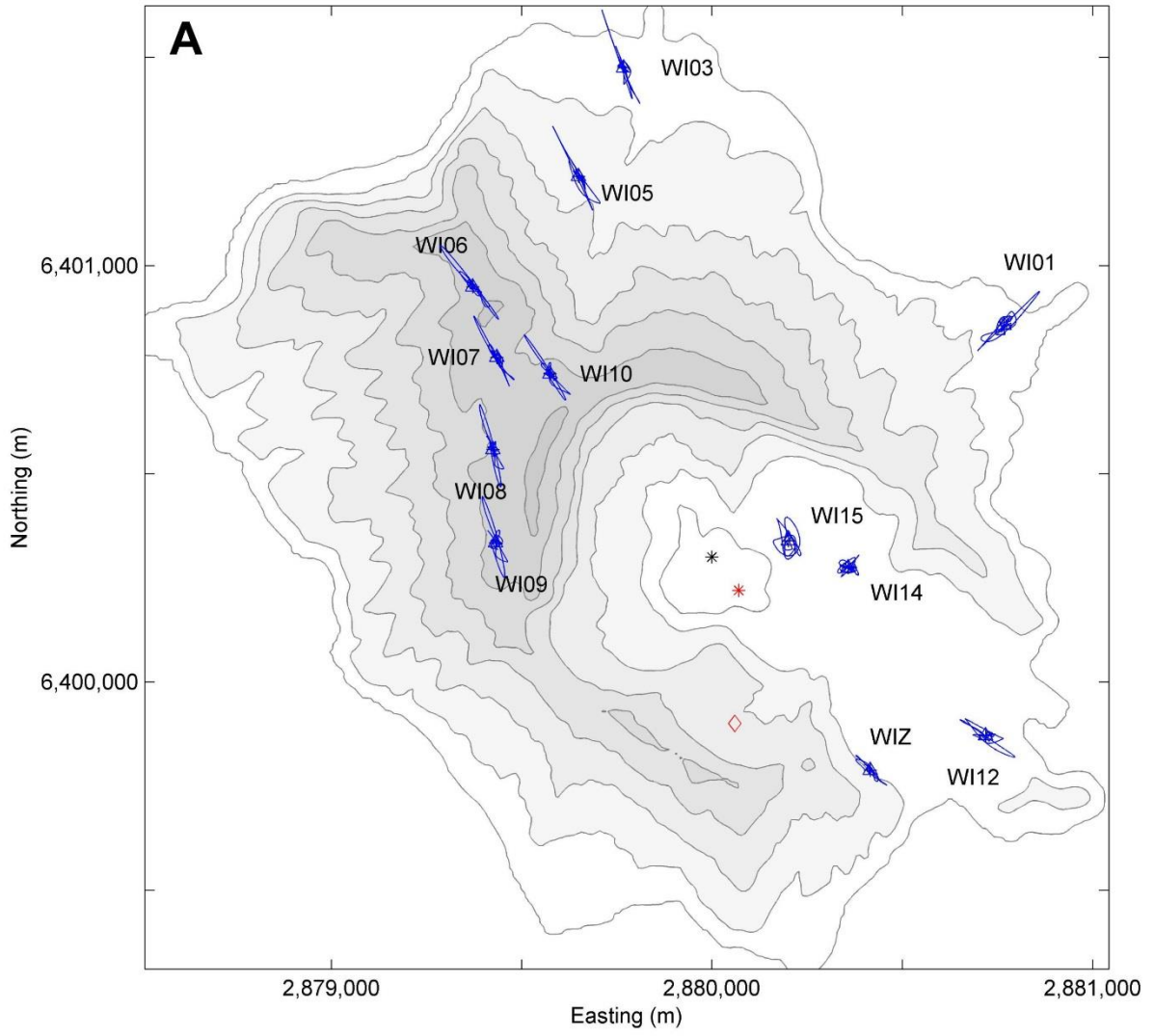
1284 Fig: 6: Sectional view of the computational model used for the numerical calculation of  
 1285 synthetic seismograms. The free surface is modelled after the topography of White Island and  
 1286 its surrounding bathymetry. A homogeneous velocity model ( $V_P = 2.2$  km/s) with otherwise  
 1287 identical model geometry was used for the synthetic test and the HF, LP, and VLP locations  
 1288 as well as the inversion analysis.

1289

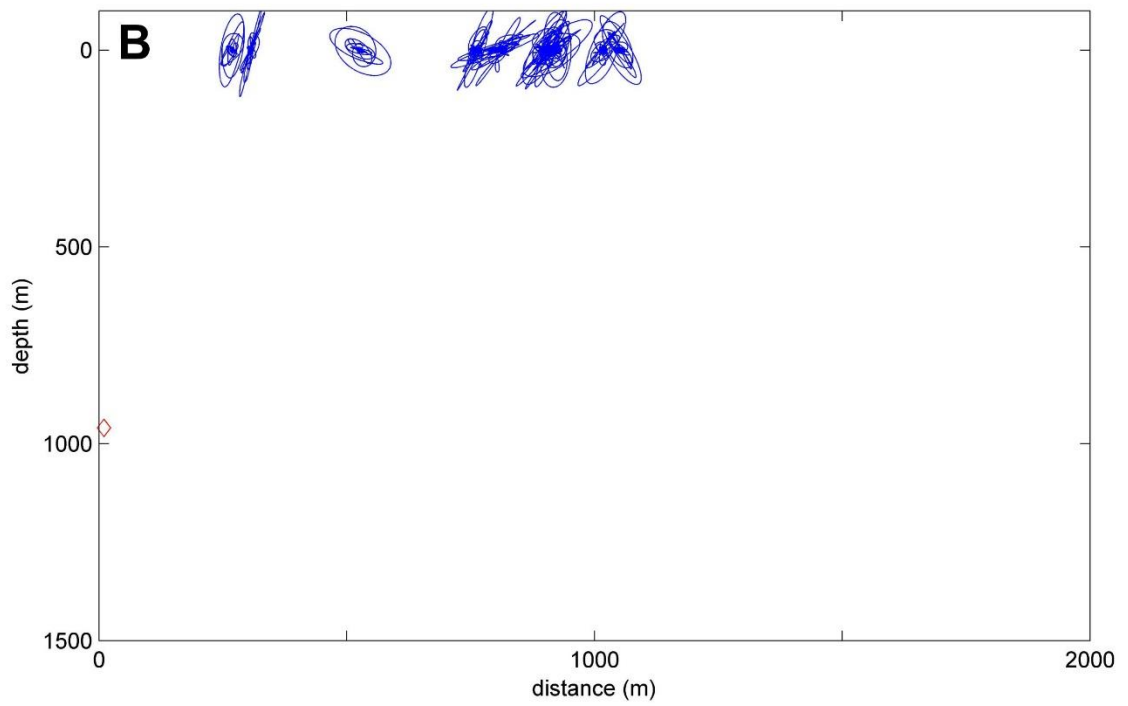
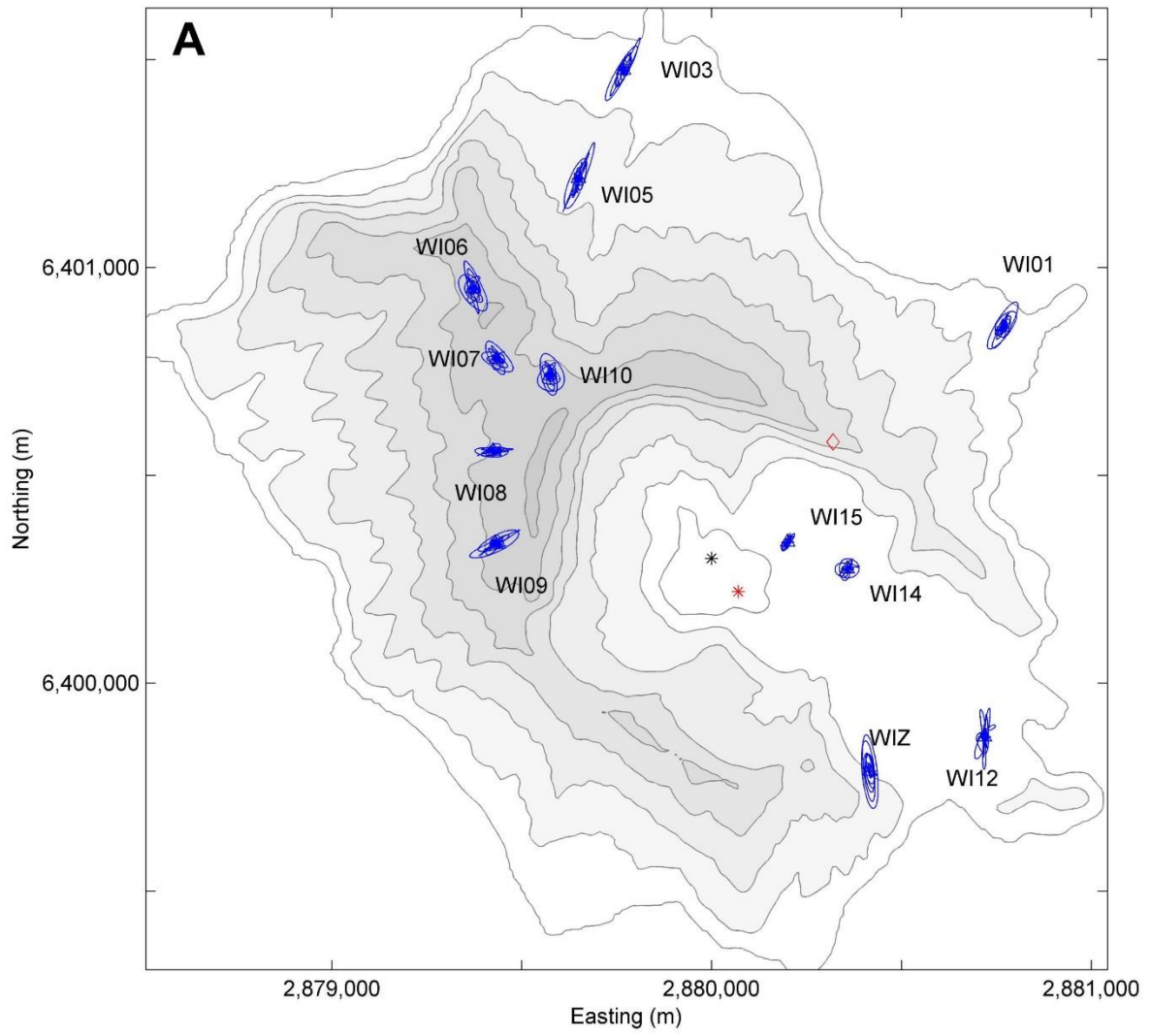
1290



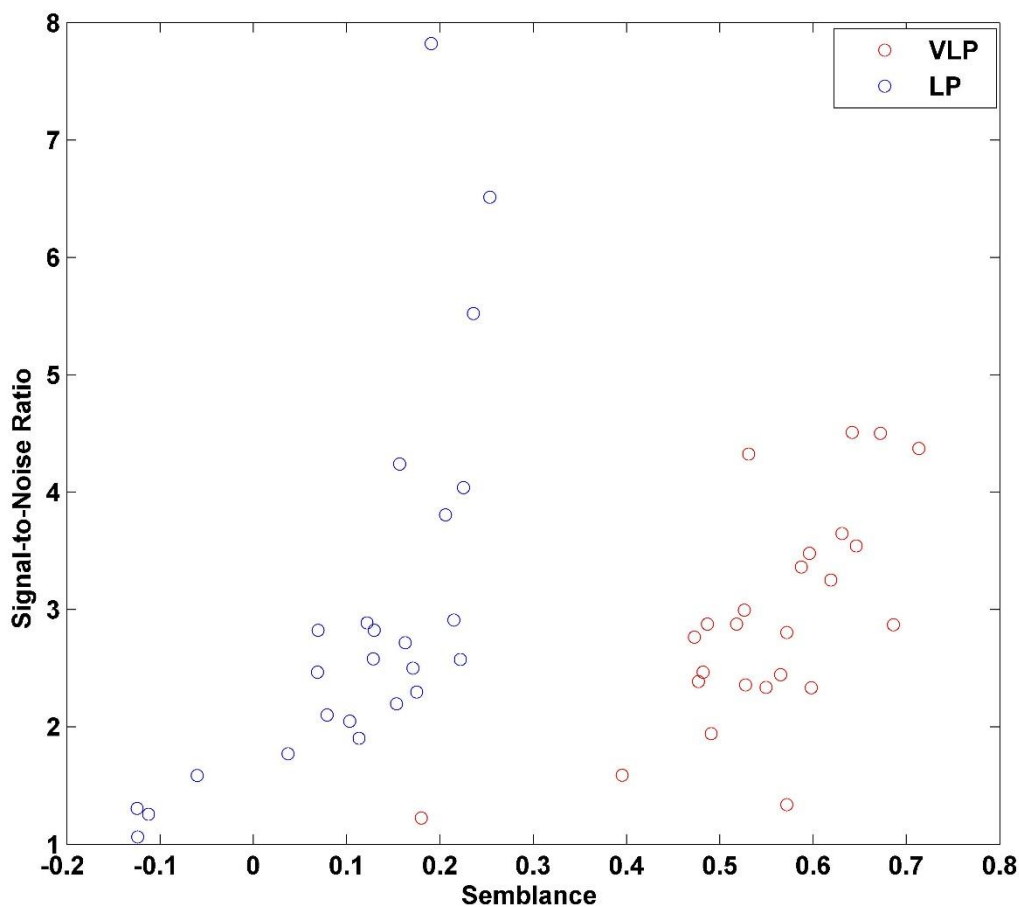
1292 Fig 7: Map view particle motions (A) for synthetic isotropic source located at 1 km below sea  
1293 level (Case 1). Synthetics are created using the spectral element method outlined in Section  
1294 4.2. The red star is a known eruption vent and the black star is a small dome emplaced in  
1295 September to November 2012, while the red circle is the location derived for the synthetic  
1296 seismograms and blue filled circle is the best fit location based on waveform semblance as  
1297 described in Kawakatsu et al., (2000). Particle motions (B) projected to vertical plain for  
1298 synthetic seismograms at the best fit location determined from waveform semblance. The  
1299 best fit location matches closely the location of the synthetic source, validating the waveform  
1300 semblance codes.



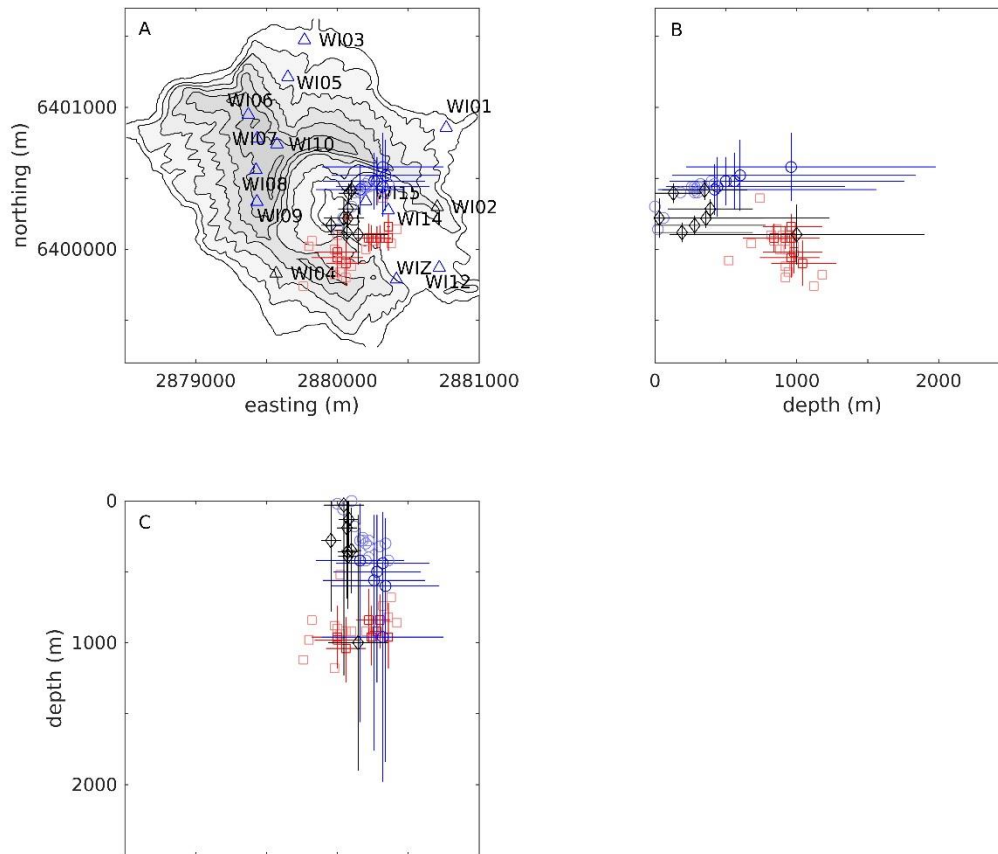
1302 Fig 8: Map view particle motions (A) for VLP component of event 17. The best fit location  
1303 is approximately 300 m south of the active vent region. Section view particle motions (B)  
1304 projected from the best fit location (red diamond) based on waveform semblance. Note the  
1305 general decrease in slant angle with distance from the inferred source location.



1307 Fig 9: Map view particle motions (A) for LP component of event 17. The best fit location  
1308 (red diamond) is approximately 300 m northeast of the active vent region. Section view  
1309 particle motions (B) projected from the best fit location determined using maximum  
1310 waveform semblance. Note the poor agreement in particle motions here interpreted as a  
1311 significant departure from an isotropic source pattern.

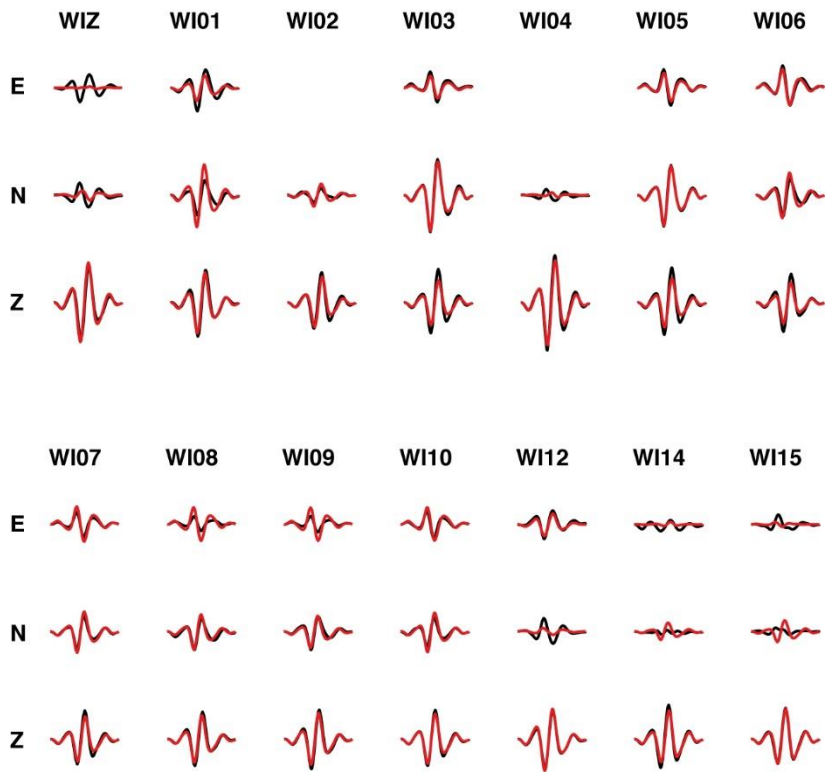


1312  
1313 Fig 10: Comparison of signal to noise ratio (SNR) and semblance (see Table 1 and 2) for 25  
1314 LP and VLP earthquakes. Signal to noise is computed in the passbands outlined in Fig 3 and  
1315 are computed for each event compared to the corresponding pre-event noise window. Note  
1316 strong VLP correlation between SNR and computed semblance for SNR below 3.



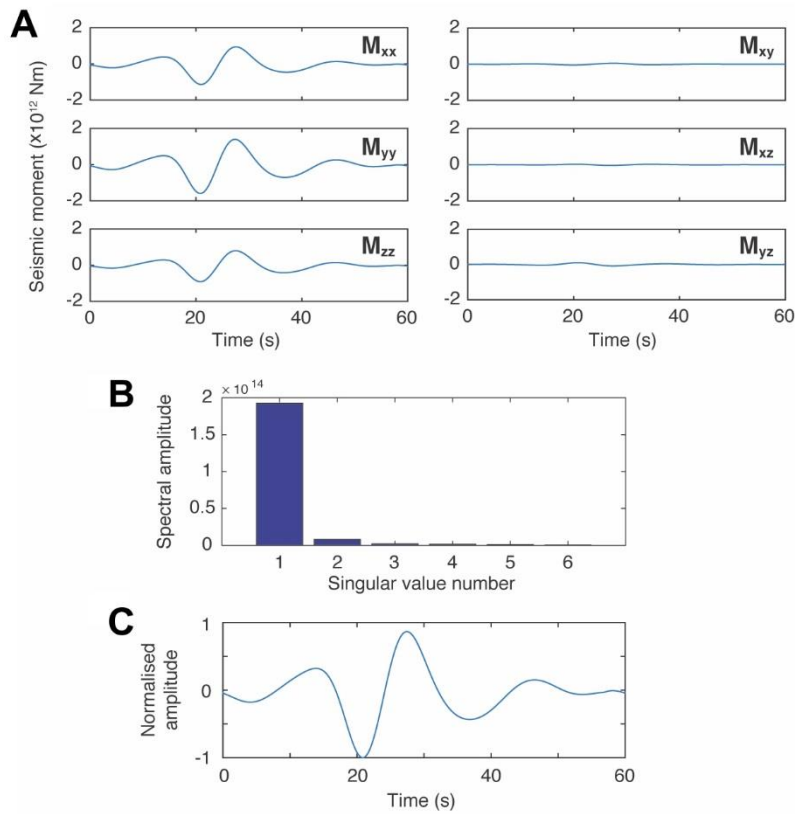
1317

1318 Fig 11: Locations and error bounds for VLP (red), LP (blue) and HF (black) sources. LP and  
 1319 VLP sources are located using waveform semblance approach outlined in Section 4.1. HF  
 1320 locations are determined using Hypo71, via triangulation of picked arrival times in the  
 1321 temporary network. See Table 1 for detailed summary of the results.



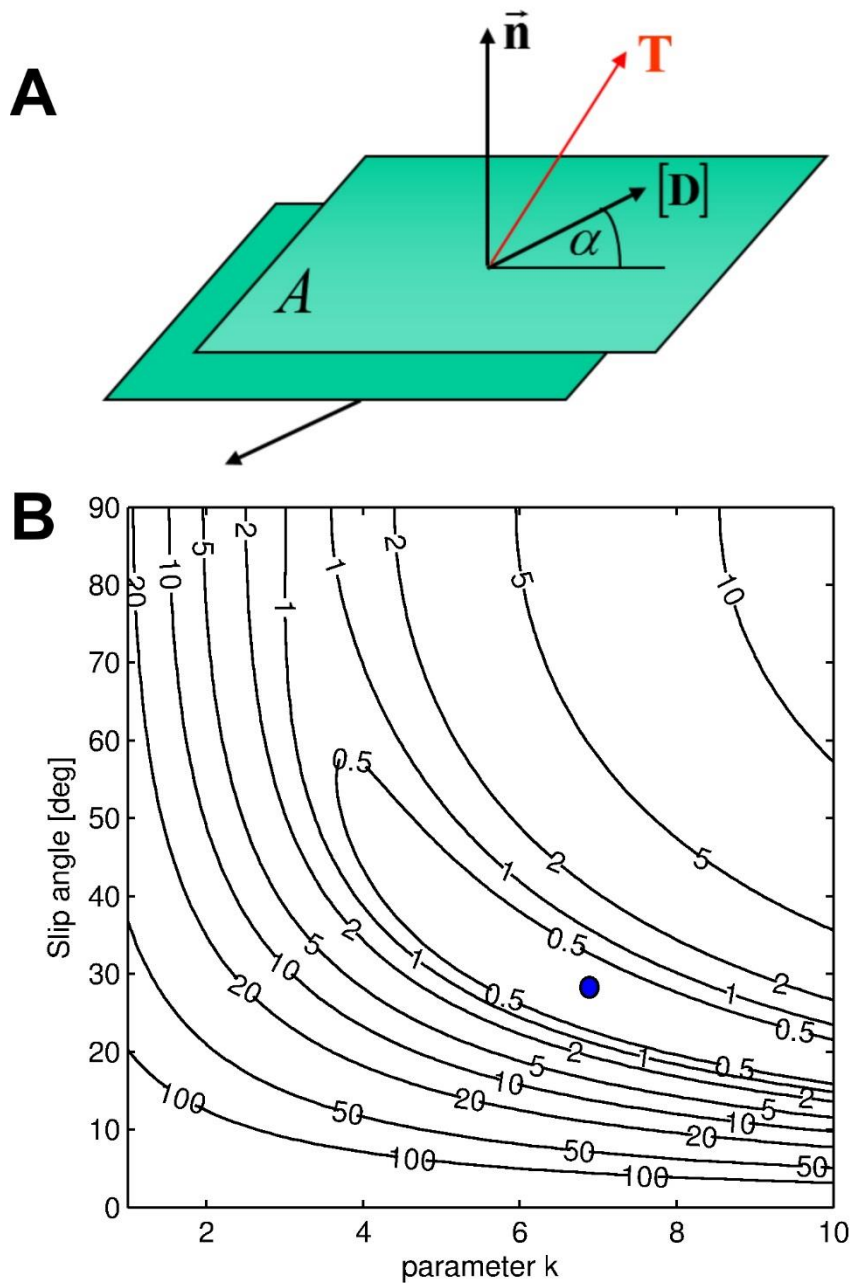
1322

1323 Fig 12 Waveform fit for the moment tensor inversion (MTI) of the VLP stack without single  
 1324 forces. The original seismograms for all stations and components are shown in black, the  
 1325 corresponding synthetic seismograms for the MTI result in red. The faulty East components  
 1326 of the stations WI02 and WI04 which were not used for the inversion are blank. A common  
 1327 scale is used for all seismograms.



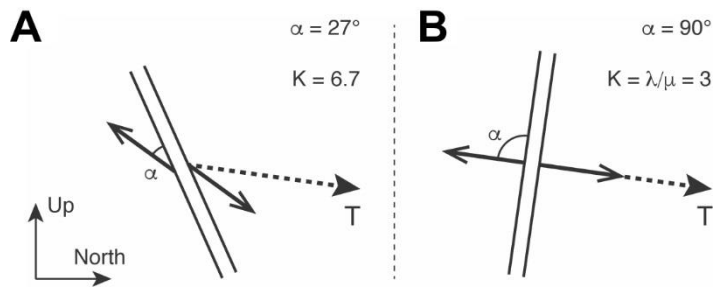
1328

1329 Fig 13 Moment tensor inversion results for the VLP stack. A) Full solutions of all 6 MT  
 1330 components, exhibiting most of the energy on the diagonal elements (left), whereas little  
 1331 energy falls on the off-diagonal components (right). B) Spectral amplitude of the singular  
 1332 values of the Vasco (1989) decomposition, showing that a unique source time function can be  
 1333 found for all MT components (shown in C)). This STF, multiplied with the MT components  
 1334 in Table 4, yields a robust representation of the MTI results.



1335

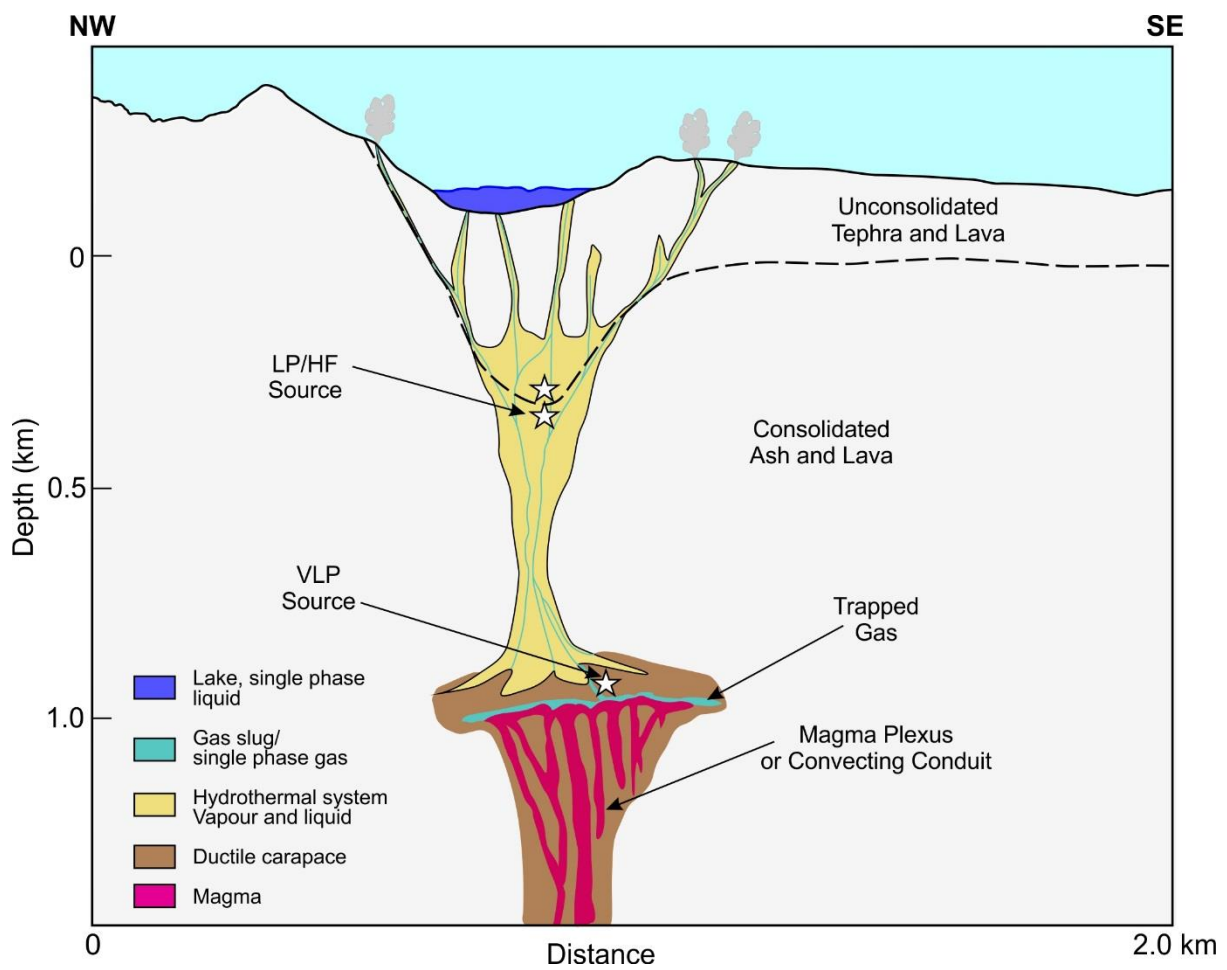
1336 Fig 14 A) A sketch of the tensile earthquake.  $\mathbf{D}$  is the slip vector on the fault and its direction  
 1337 is determined by the angle  $\alpha$  from the fault plane. B) Contour plot of  $R_2$  misfit (in percentage)  
 1338 between the theoretical and observed principal values of moment-tensor for a general non-DC  
 1339 crack source. The blue solid circle denotes the minimum misfit. However, the whole region is  
 1340 within the 1% misfit range and can be considered as viable candidates for the source  
 1341 mechanism.



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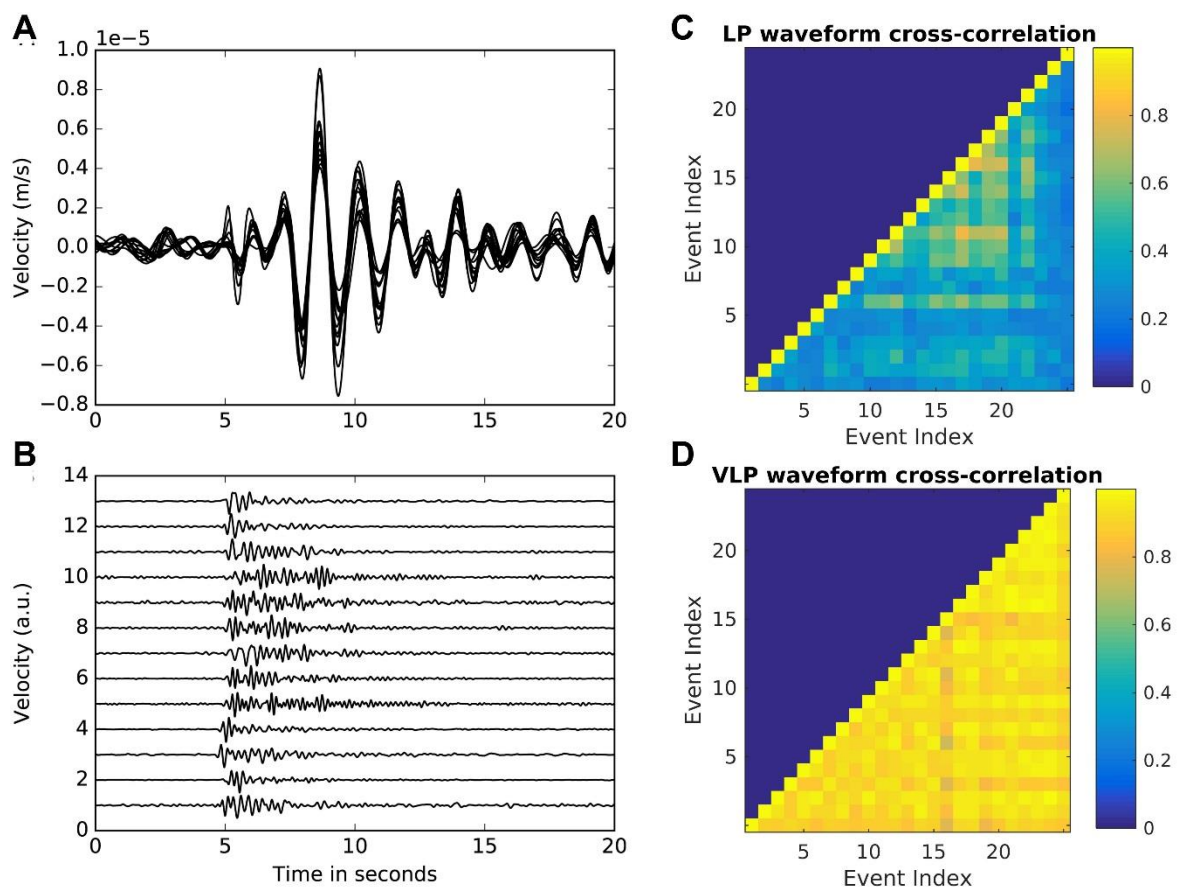
1343 Fig 15. Illustration of the VLP source mechanism. A) The best solution suggests a quasi-  
 1344 vertical tensile crack with a significant amount of shear movement. B) An alternative  
 1345 solution, also with a relatively low residual. The examples illustrate our inability to robustly  
 1346 resolve between the different mechanisms.

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1349 Fig 16 Conceptual model of coupled VLP/LP/HF source processes. VLP source is inferred to  
 1350 be at the top of the magma carapace at ~1 km depth. Gas slug release propagates to shallow  
 1351 depths inducing failure of unconsolidated ash and lava materials in the vent or vent walls  
 1352 (Heap et al, 2016). Alternatively, the shallow hydrothermal/vent fill system may include  
 1353 strong cavity resonance characteristics that may generate LP (e.g., Chouet 1996; Neuberg et  
 1354 al, 2002). Localised HF seismicity may result from failure due to stress concentration near to  
 1355 a localised conduit constriction.



1356

1357 Fig 17: A) Vertical LP waveforms (event 11, filtered between 0.4 and 1.2 Hz) aligned for all  
 1358 stations maximising the cross-correlation between all trace pairs. Station WI04 was omitted  
 1359 due to a considerably different waveform. B) The same traces filtered between 0.4 and 5 Hz  
 1360 and plotted individually. Note that the HF waveforms are incoherent between stations and  
 1361 their onset time relative to the LP vary by a few tenths of a second. This suggests that HF

1362 and LP signals originate at slightly different locations. C) waveform cross-correlation matrix

1363 for LP earthquakes, D) equivalent cross-correlation matrix for VLP events.

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